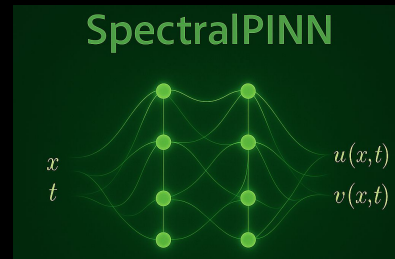




# Teukolsky by Design: A Hybrid Spectral-PINN solver for Kerr Quasinormal Modes

Alexandre M. Pombo



## Introduction: Quasinormal modes

- Gravitational wave detections are a window to explore compact objects
- These are ideal probes of fundamental and dark-sector physics
- Quasi-normal modes encode stability and the ringdown information
- The resulting GW carries information about the merging objects and the final remnant

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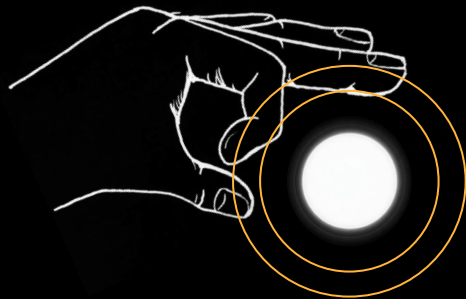
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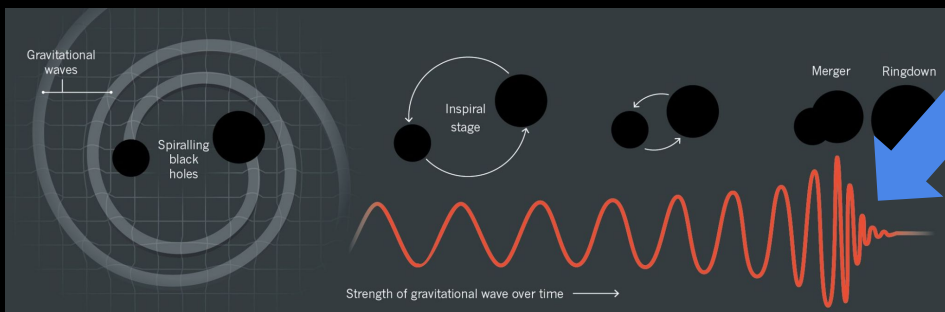
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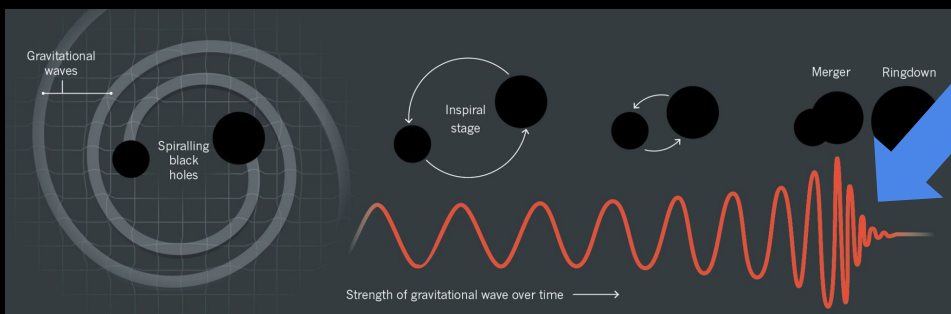
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- QNM computation is challenging, leaving many models unexplored.

## Kerr BH: Teukolsky equation

$$\begin{aligned} & \left[ \frac{(r^2 + a^2)^2}{\Delta} - a^2 \sin^2 \theta \right] \frac{\partial^2 \psi}{\partial t^2} + \frac{4Mar}{\Delta} \frac{\partial^2 \psi}{\partial t \partial \varphi} + \left[ \frac{a^2}{\Delta} - \frac{1}{\sin^2 \theta} \right] \frac{\partial^2 \psi}{\partial \varphi^2} \\ & - \Delta^{-s} \frac{\partial}{\partial r} \left( \Delta^{s+1} \frac{\partial \psi}{\partial r} \right) - \frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial \psi}{\partial \theta} \right) - 2s \left[ \frac{a(r - M)}{\Delta} + \frac{i \cos \theta}{\sin^2 \theta} \right] \frac{\partial \psi}{\partial \varphi} \\ & - 2s \left[ \frac{M(r^2 - a^2)}{\Delta} - r - ia \cos \theta \right] \frac{\partial \psi}{\partial t} + (s^2 \cot^2 \theta - s) \psi = 0, \end{aligned}$$



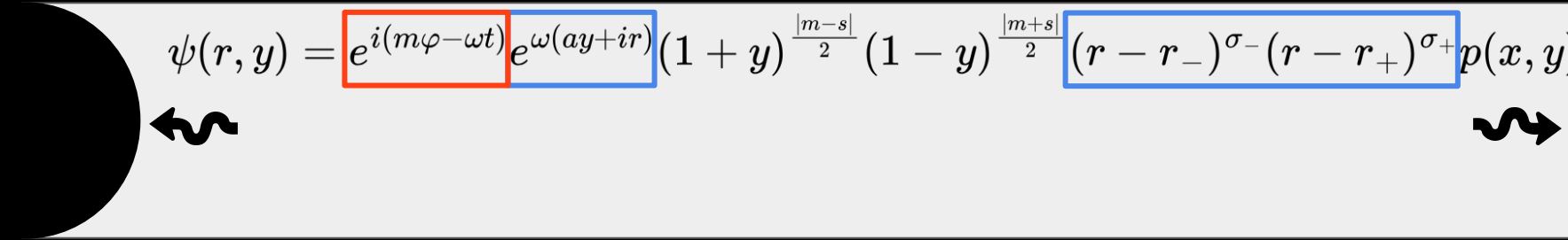
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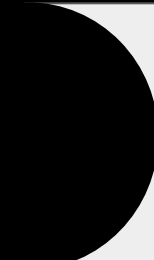
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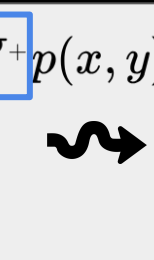


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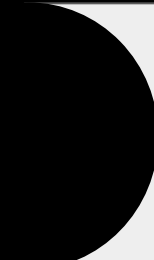
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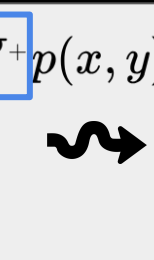


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$$p(x, y) = f(x)g(y)$$

## Normalization conditions

- The Teukolsky QNM are a homogeneous linear eigenvalue problem
- QNM amplitudes are fixed by the initial data rather
- Once  $(\omega, \Lambda)$  is fixed the equations admit:  $(f, g) \rightarrow (C f, C^{-1} g) \quad \square \quad p \rightarrow C p$
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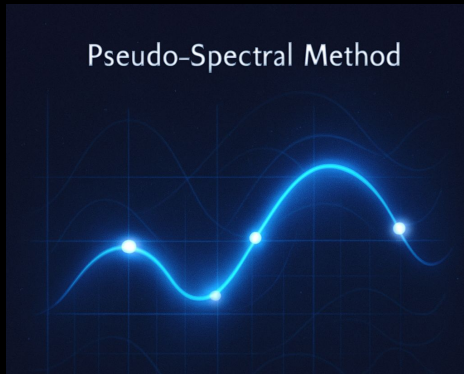
$$f(x) = (e^{x-1} - 1)\mathcal{F}(x) + 1$$

$$g(y) = (e^{y+1} - 1)\mathcal{G}(y) + 1$$

$$p(x, y) = 1 + [\mathcal{P}(x, y) - \mathcal{P}(1, -1)]$$

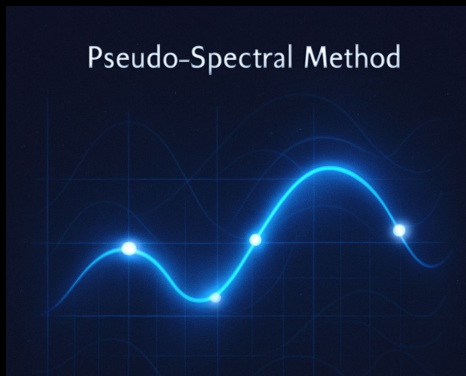
# Pseudo-Spectral Method

$$p(x,y) = \sum A_{i,j} T_i(x) T_j(y)$$

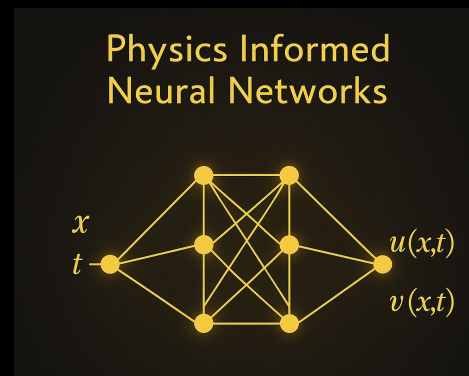
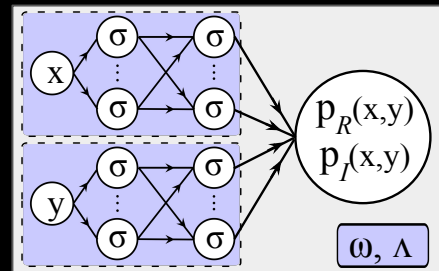


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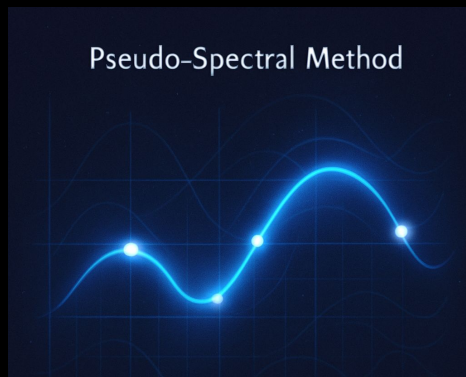


## Physics Informed Neural Networks

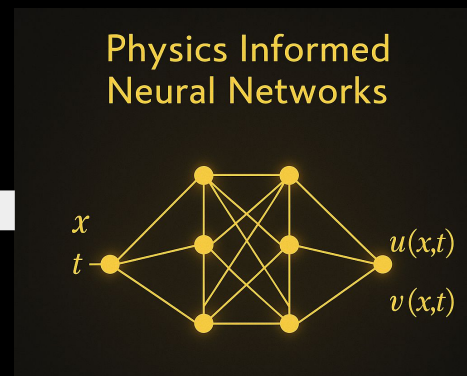
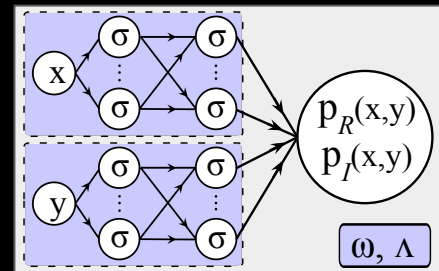


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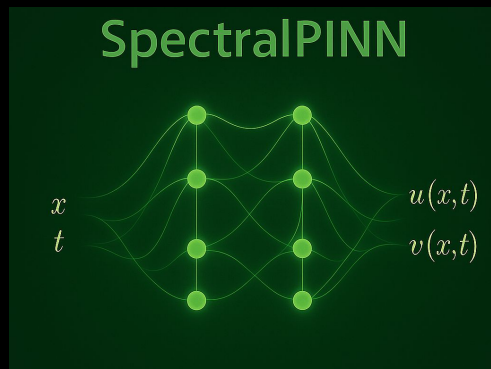


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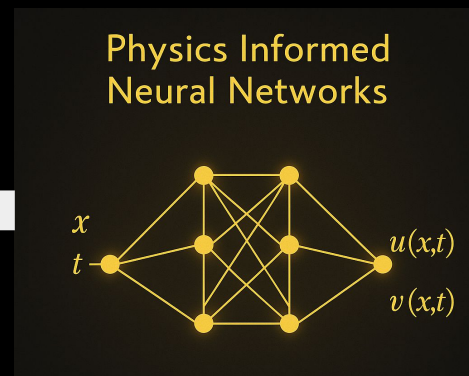
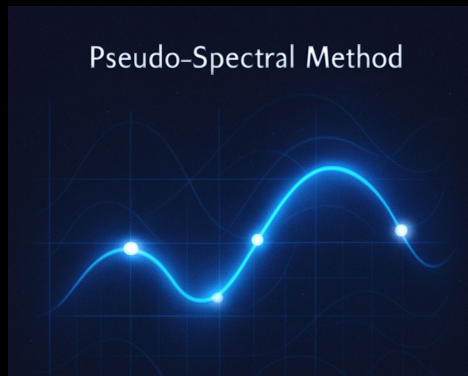
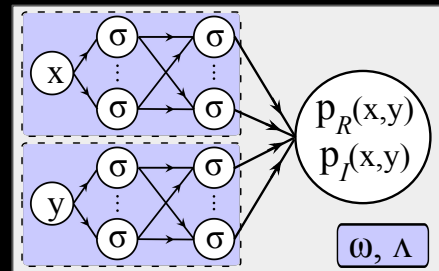


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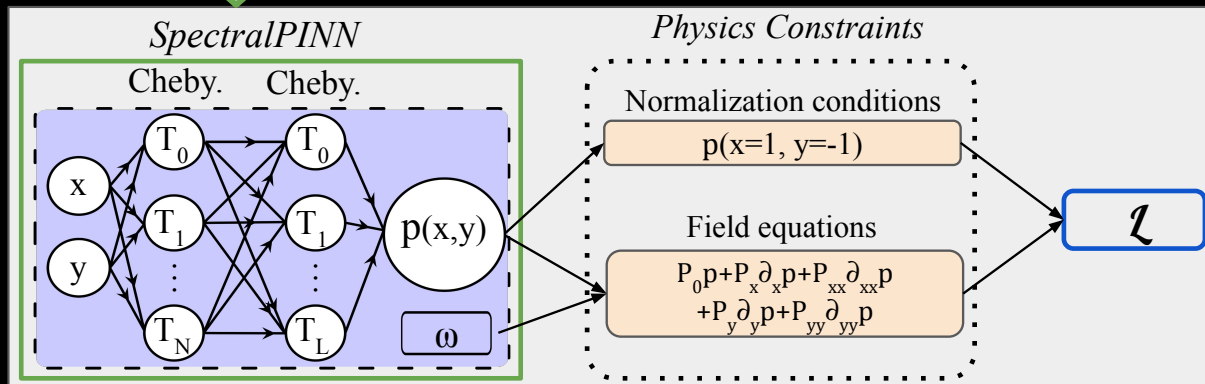
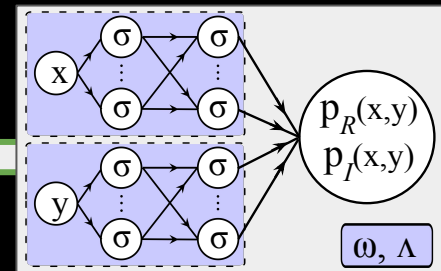


# SpectralPINN

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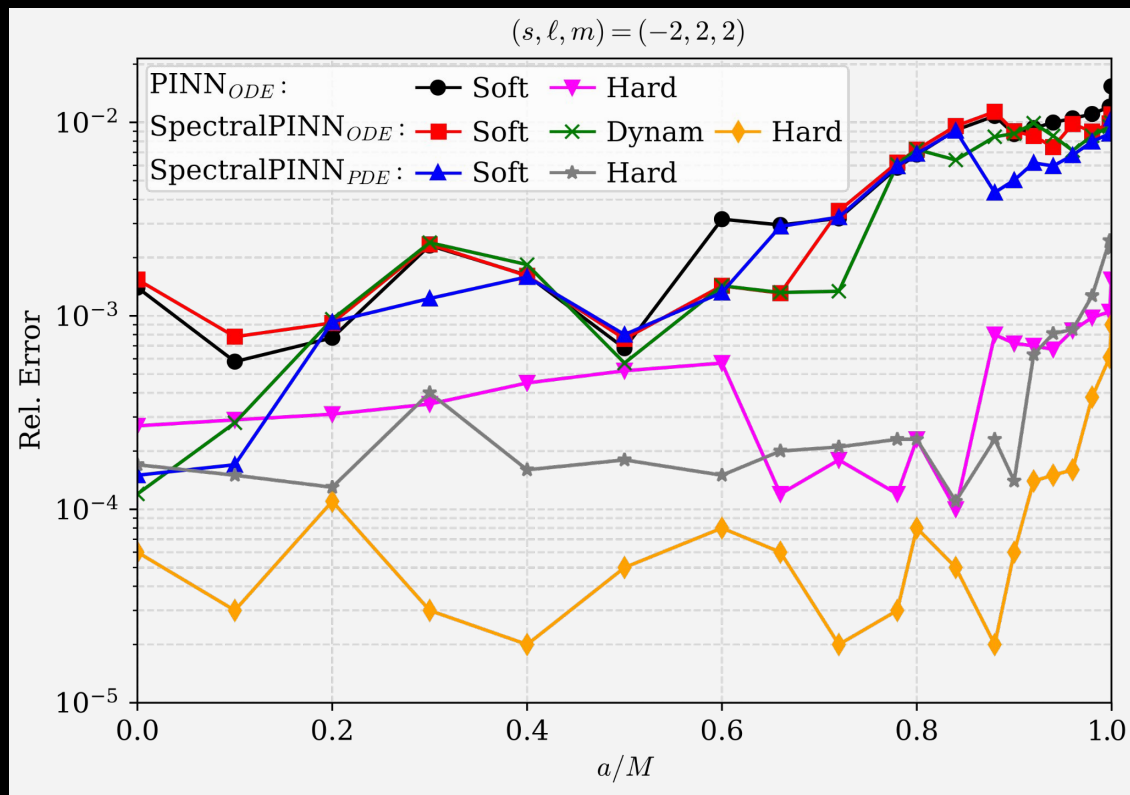
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# Results

## Results: Numerical accuracy



## Results: Quadrupolar deviation

- The SpectralPINN formulation is able to solve the 2D PDE
- Perturbations of Kerr that spoil the separability of Teukolsky equation
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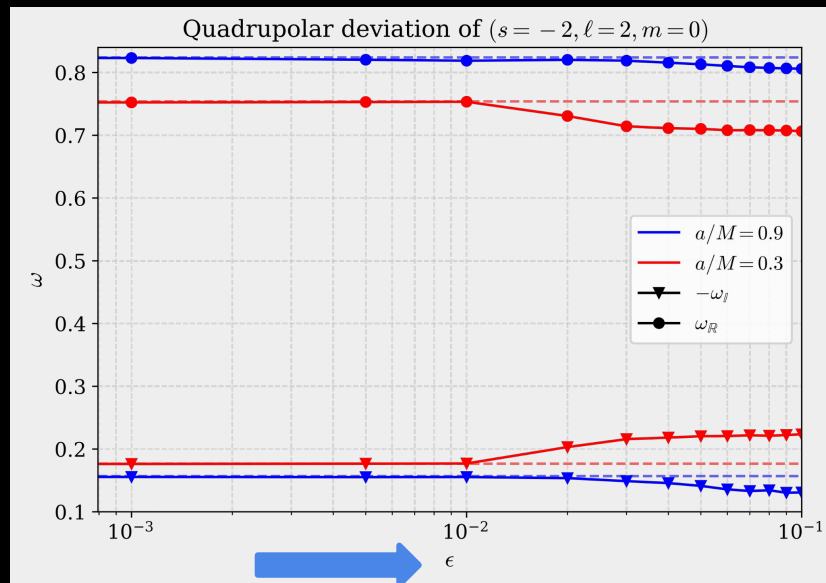
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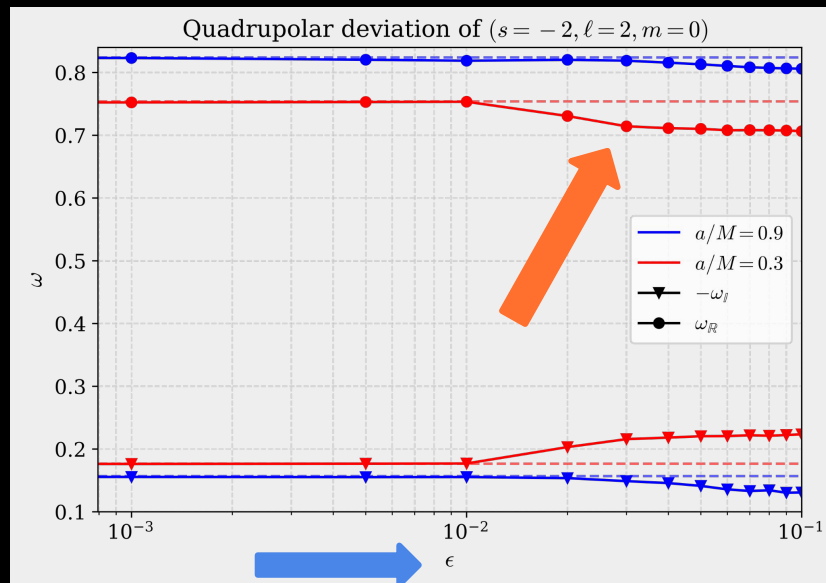


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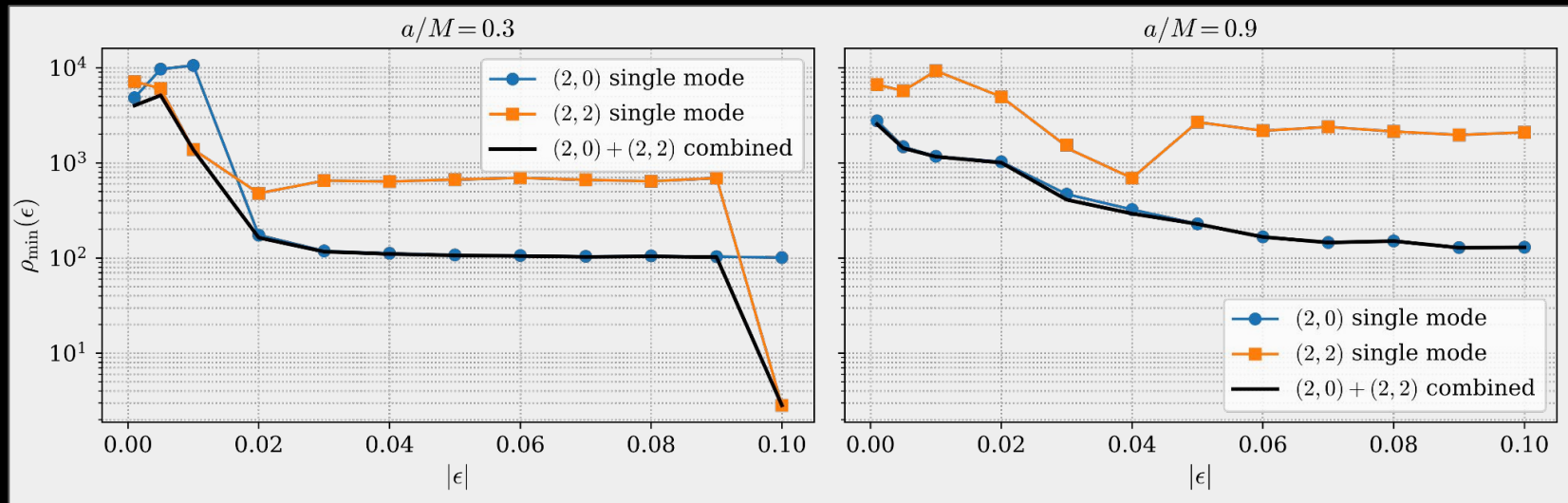
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# Results: Einstein Telescope



# Conclusion

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Quasi-Normal  
Modes

SpectralPINN



Recent and  
near future  
observations

Obrigado !

Thanks !



# Teukolsky by Design: A Hybrid Spectral-PINN solver for Kerr Quasinormal Modes

Thank you!

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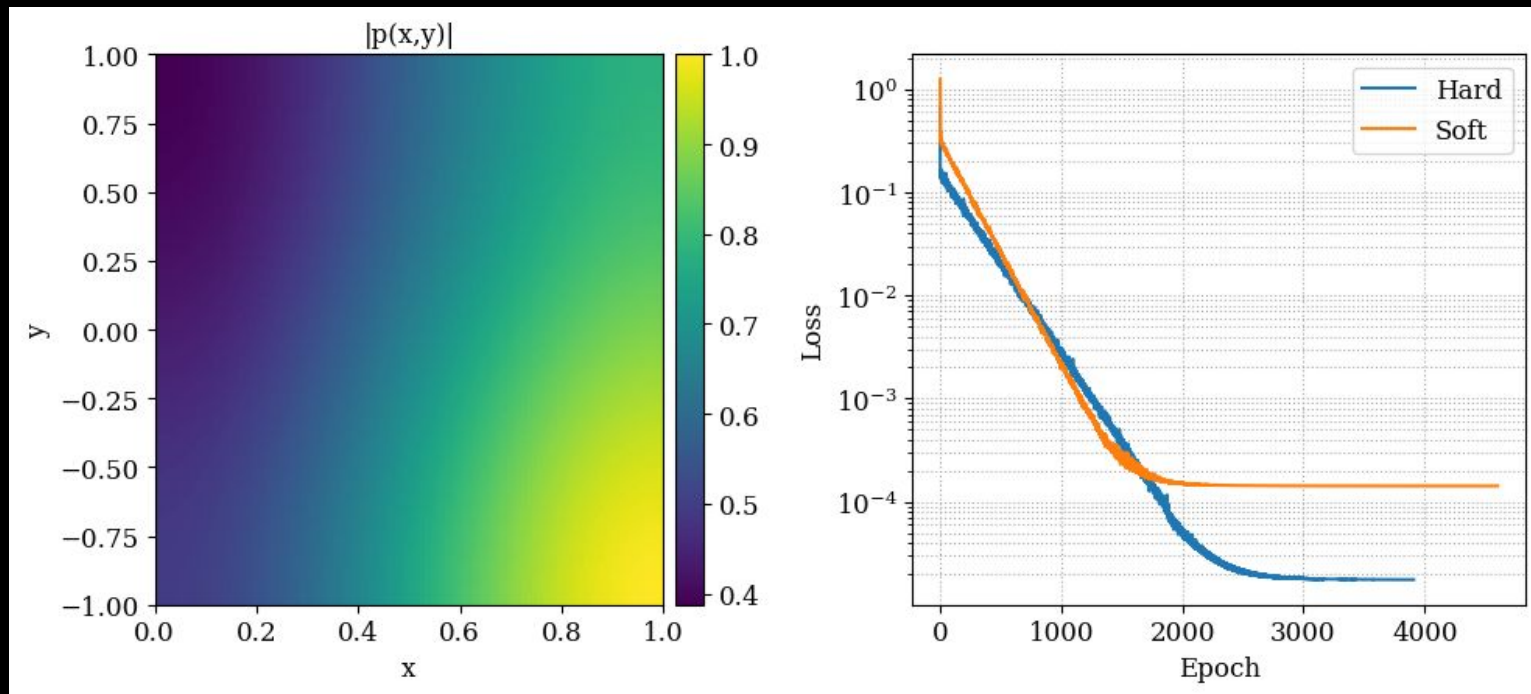
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## Extra quirks: Training



## Extra quirks: Basis number

Eq. type	$N \times L$						
	$10 \times 10$	$15 \times 10$	$10 \times 15$	$15 \times 15$	$20 \times 20$	$30 \times 30$	$50 \times 50$
ODE	0.714	0.109	0.156	0.012	0.006	0.002	0.429
PDE	0.920	0.234	0.280	0.071	0.052	0.011	0.600

## Extra quirks: Training improvements

Scheme	Eq. type	Base	+ $\mathbb{C}$ -dtype	+HAdamD	+Alt. Sched.
PINN	ODE	0.138	0.113	0.069	0.050
SpectralPINN	ODE	0.131	0.126	0.016	0.002
SpectralPINN	PDE	0.142	0.134	0.058	0.011