

# Backreaction effects in a recursive Penrose process in Reissner-Nordström-AdS black hole spacetimes

**University of Helsinki**



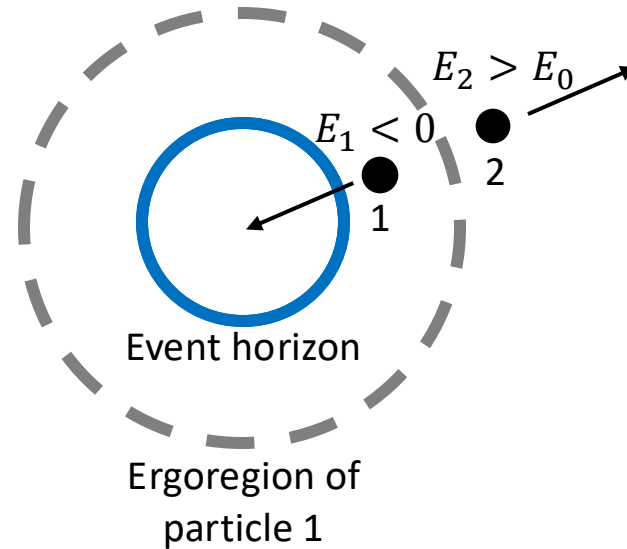
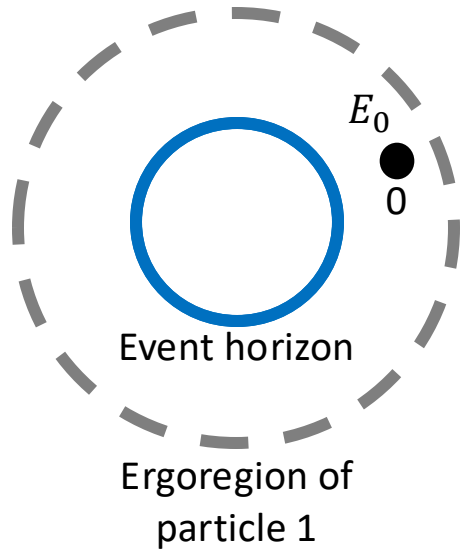
Duarte Feiteira

In collaboration with: Prof. José P. S. Lemos, Prof. Oleg B. Zaslavskii



Feiteira, Lemos, Zaslavskii; Phys.Rev.D 109 (2024) 6, 064065

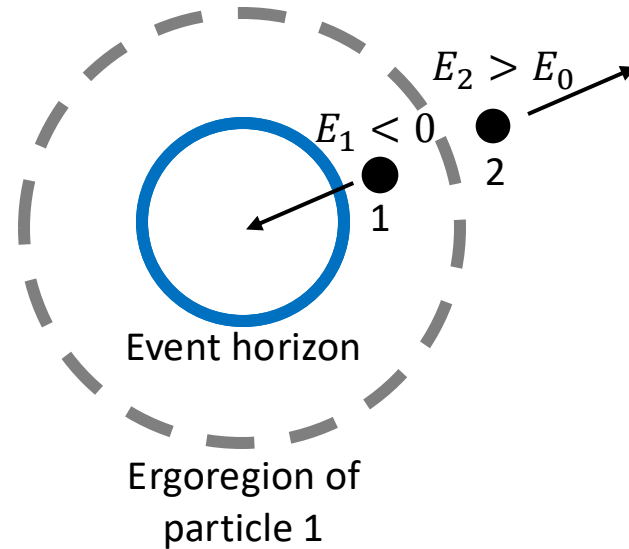
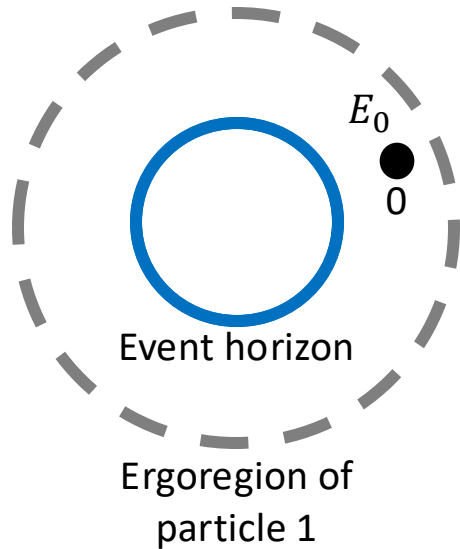
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- **Penrose process:**



- Particle 2 is reflected: **recursive Penrose process**.
- A recursive Penrose process has been studied in Reissner-Nordström-AdS black hole spacetimes **neglecting backreaction effects**.
- Two possible situations: Black hole **energy factories** and black hole **bombs**.
- A infinitely long chain of decays is expected to **change significantly the charge and mass of the black hole**.

- Line element of RN-AdS black hole spacetime:

$$ds^2 = -f(r) dt^2 + \frac{dr^2}{f(r)} + r^2 d\Omega^2, \quad f(r) = \frac{r^2}{l^2} + 1 - \frac{2M}{r} + \frac{Q^2}{r^2}$$

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- Equations of motion for particles with mass  $m$  and charge  $e$ :

$$p^t \equiv m \dot{t} = \frac{X}{f}, \quad p^r \equiv m \dot{r} = \sigma P$$

$$X = E - \frac{eQ}{r}, \quad P = \sqrt{X^2 - m^2 f}$$

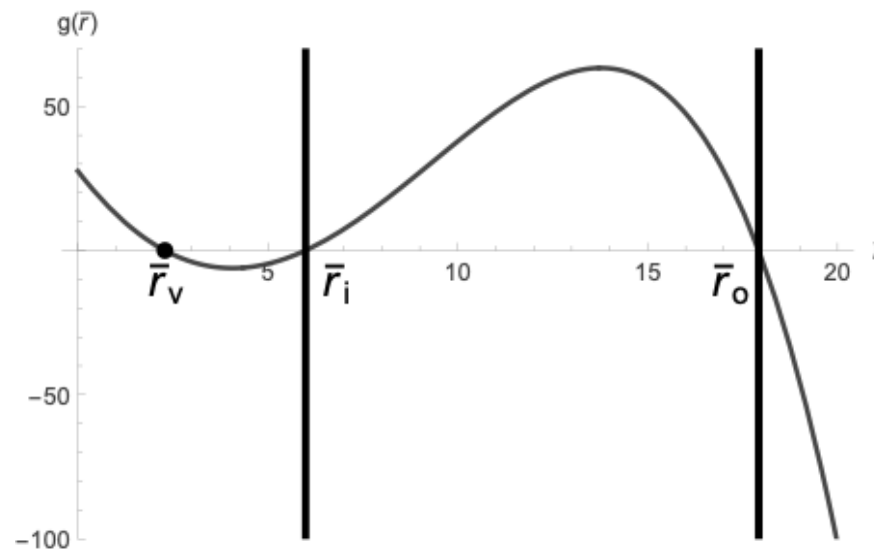
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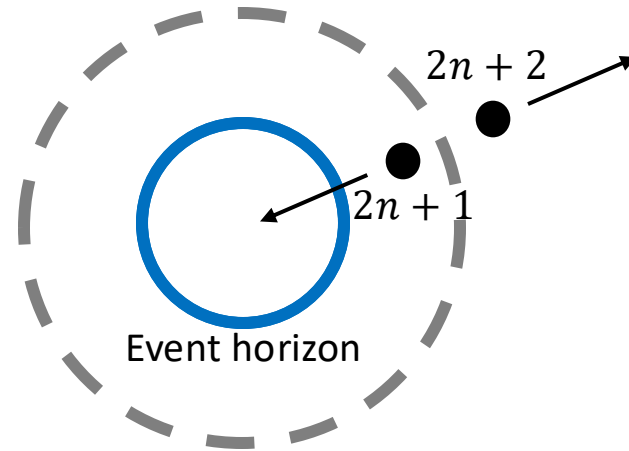
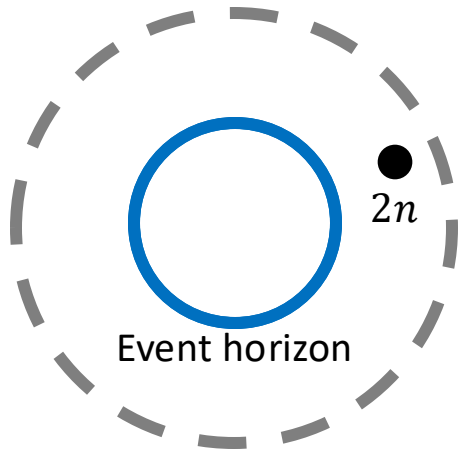
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- Just after the  $(n + 1)$ -th decay and right before particle  $2n + 1$  falls into the black hole:

$$M_n = M_0 + E_1 + E_3 + \cdots + E_{2n+1} = M_0 + \sum_{k=0}^{n-1} E_{2k+1}$$

$$Q_n = Q_0 + e_1 + e_3 + \cdots + e_{2n+1} = Q_0 + \sum_{k=0}^{n-1} e_{2k+1}$$

$$f_n(r) = \frac{r^2}{l^2} + 1 - \frac{2M_n}{r} + \frac{Q_n^2}{r^2}$$



- Assuming particles are at rest immediately after each decay,  $P = 0$ :

$$E_{2n} = m_{2n}\sqrt{f_n(r_i)} + \frac{e_{2n}Q_n}{r_i}, \quad E_{2n+1} = m_{2n+1}\sqrt{f_n(r_i)} + \frac{e_{2n+1}Q_n}{r_i}, \quad E_{2n+2} = m_{2n+2}\sqrt{f_n(r_i)} + \frac{e_{2n+2}Q_n}{r_i}$$

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- Mass of the particles:

$$m_{2n+1} = \alpha_1 m_{2n}, \quad m_{2n+2} = \alpha_2 m_{2n}$$

- Mass conservation at each decay:  $\alpha_1 + \alpha_2 = 1$

$$m_{2n+1} = \alpha_1 (1 - \alpha_1)^n m_0, \quad m_{2n+2} = \alpha_2^{n+1} m_0$$

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- Charge of the particles:

$$e_{2n+1} = \beta_1 e_{2n}, \quad e_{2n+2} = \beta_2 e_{2n}$$

- Charge conservation at each decay:  $\beta_1 + \beta_2 = 1$

- Odd particles must have negative energy:  $\beta_1 < 0$  (assuming  $Q_0 > 0$ )

$$e_{2n+1} = \beta_1 (1 - \beta_1)^n e_0, \quad e_{2n+2} = \beta_2^{n+1} e_0$$

- Charge of the black hole:

$$\left. \begin{aligned} Q_n &= Q_0 + \sum_{k=0}^{n-1} e_{2k+1} \\ e_{2n+1} &= \beta_1 (1 - \beta_1)^n e_0 \end{aligned} \right\} \quad Q_n = Q_0 - e_0 ((1 - \beta_1)^n - 1)$$

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- Case 1, integer  $n_c$ :** the charge of the black hole reaches exactly zero. The recursive Penrose process terminates naturally.
- Case 2, non-integer  $n_c$ :** the charge of the black hole approaches but does not reach zero. The last viable decay occurs for the closest integer lower than  $n_c$ . After that, spherical symmetry is no longer a valid approximation.

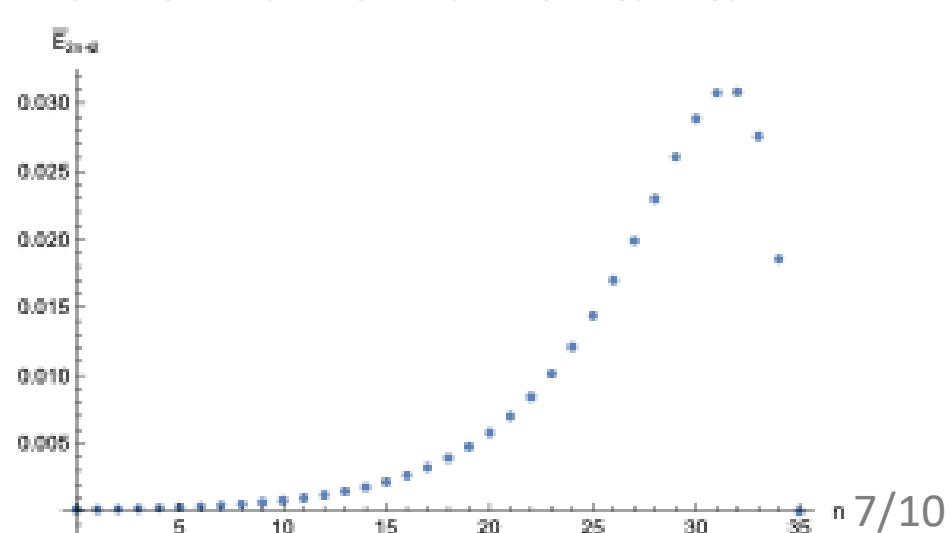
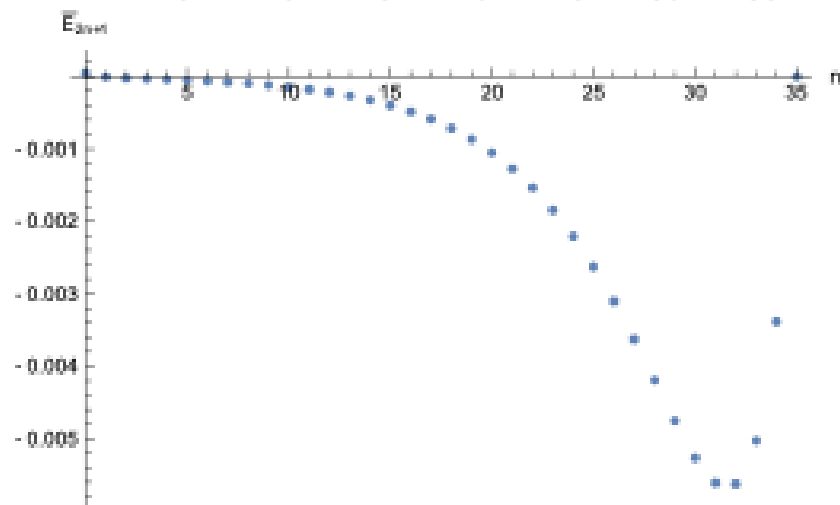
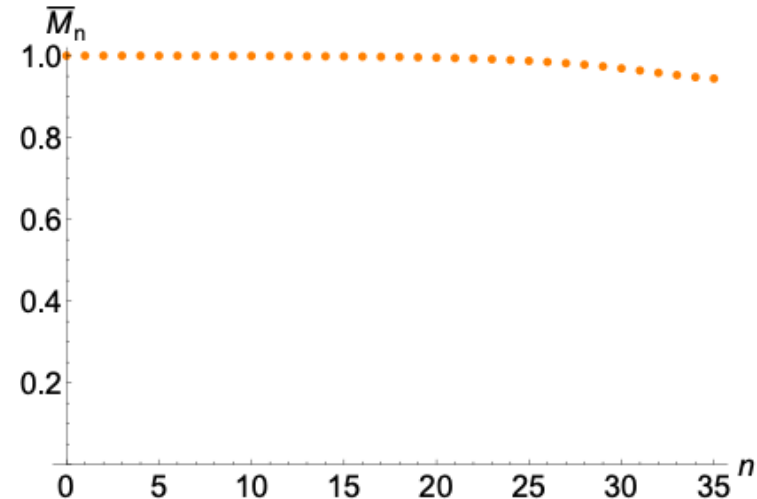
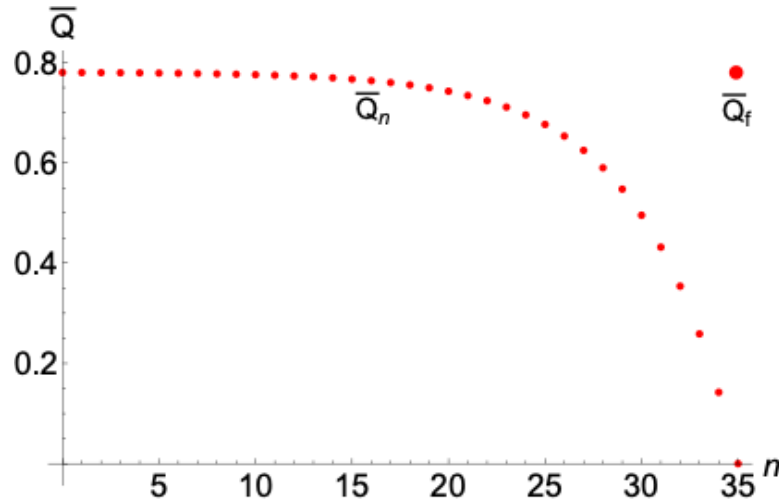
# Case 1, integer $n_c$

- Result of the process:**

Black hole with charge  $Q_0 + e_0$ ;

Finite energy gain – black hole energy factory;

No black hole bomb as obtained without backreaction effects.



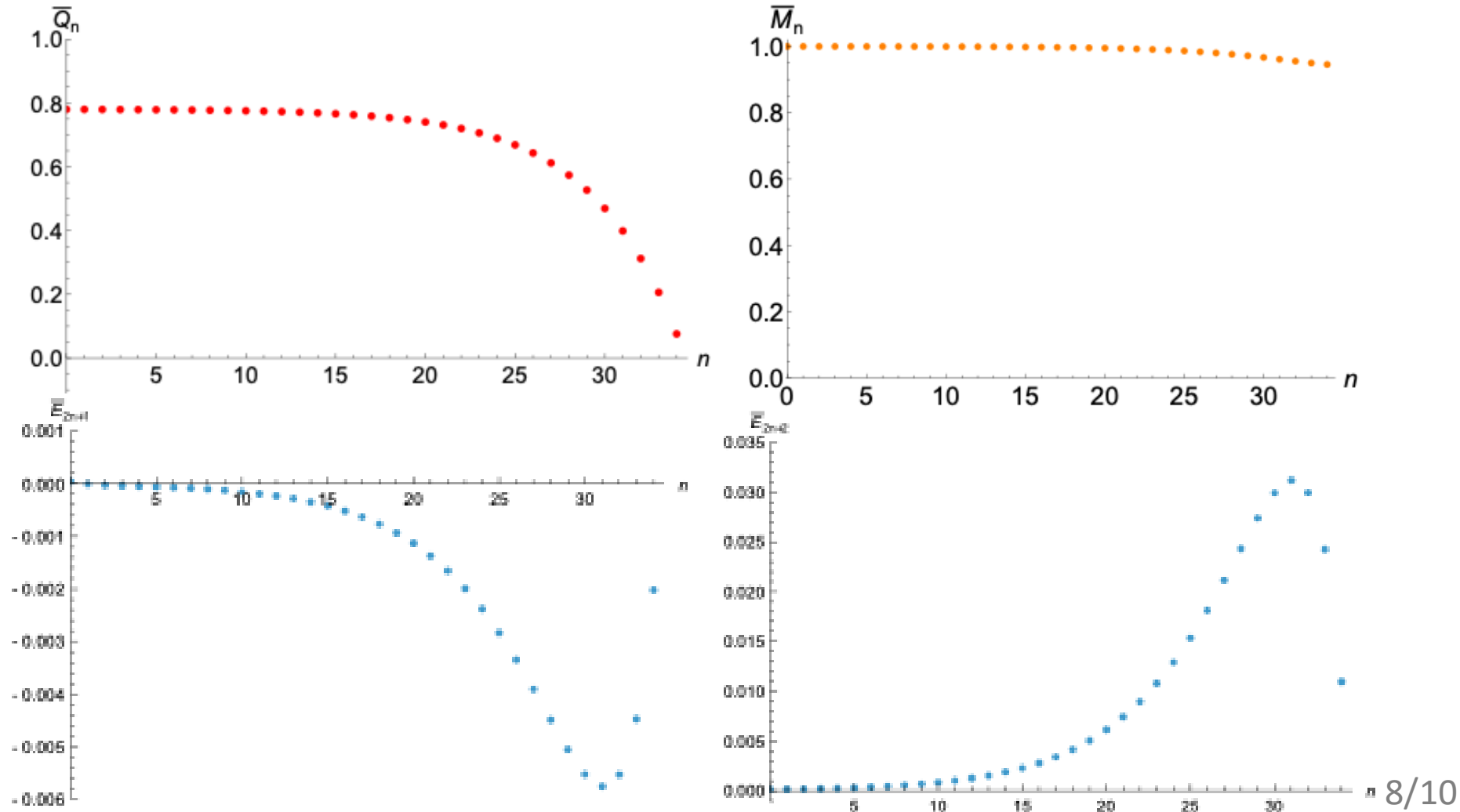
# Case 2, non-integer $n_c$

- Result of the process:**

Black hole with arbitrary small charge and particle with large charge;

No spherical symmetry – two body problem in GR;

Electric repulsion should make the two bodies fall apart.





- Backreaction effects turn black hole bombs into **black hole energy factories**.
- Particle confinement can be ensured due to the **negative cosmological constant of AdS** spacetime.
- The AdS turning point is totally **equivalent to a reflective mirror**.

- Backreaction effects turn black hole bombs into **black hole energy factories**.
- Particle confinement can be ensured due to the **negative cosmological constant of AdS** spacetime.
- The AdS turning point is totally **equivalent to a reflective mirror**.
- Depending on the initial charges of the black hole and decaying particles, one can either get at the end of the process a **black hole with very large or very small charge**.
- For **non-integer index  $n_c$** , one cannot resolve the full process within a spherical symmetry and test particle approximation: **2-body problem in GR**.

Thank you!



*Oulanka National Park, Finland*