

From Stars to Regular Black Holes

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Based on 2505.09680 & 2512.19796

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Regular Black Holes from Pure Gravity

[Bueno, Cano, Hennigar '24]

$$S = \int \left(\mathcal{L}_{\text{GR}} + \sum_{n=2}^{\infty} \alpha_n \mathcal{L}_{(n)} \right)$$

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$$ds^2 = -N(r)^2 f(r) dt^2 + \frac{dr^2}{f(r)} + r^2 d\Omega^2$$

A Regular Black Hole: Hayward

[Hayward '05]

$$\alpha_n = \alpha^{n-1}$$

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[Hayward '05]

$$\alpha_n = \alpha^{n-1} \quad \Rightarrow \quad f(r) = 1 - \frac{2Mr^2}{r^{D-1} + 2M\alpha}$$

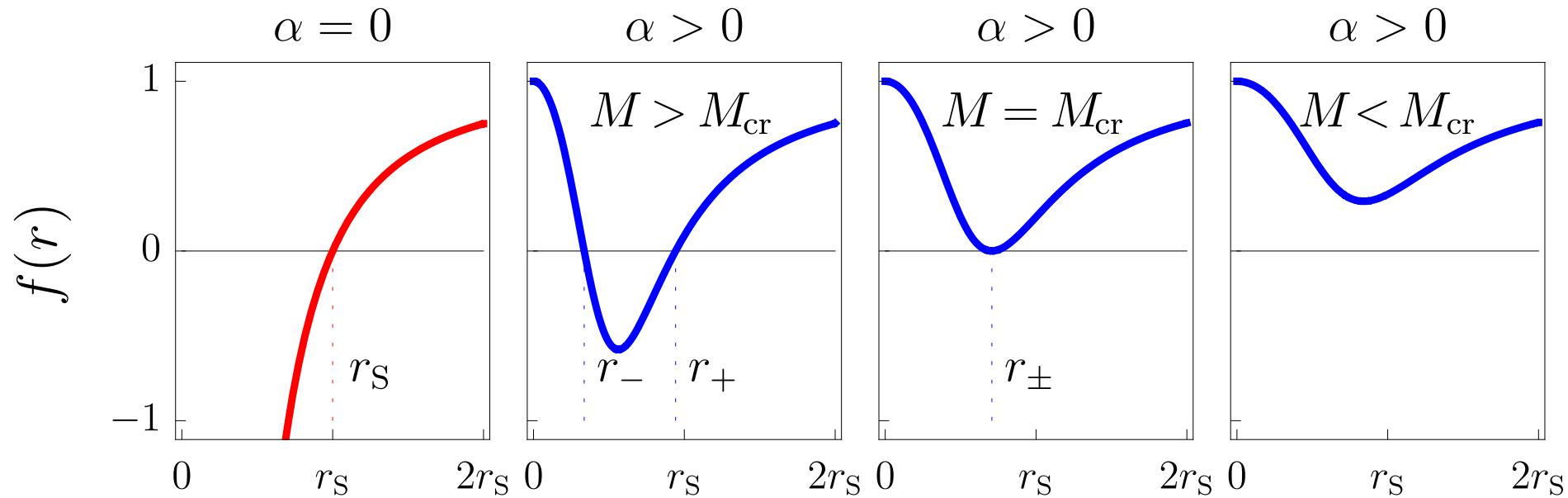
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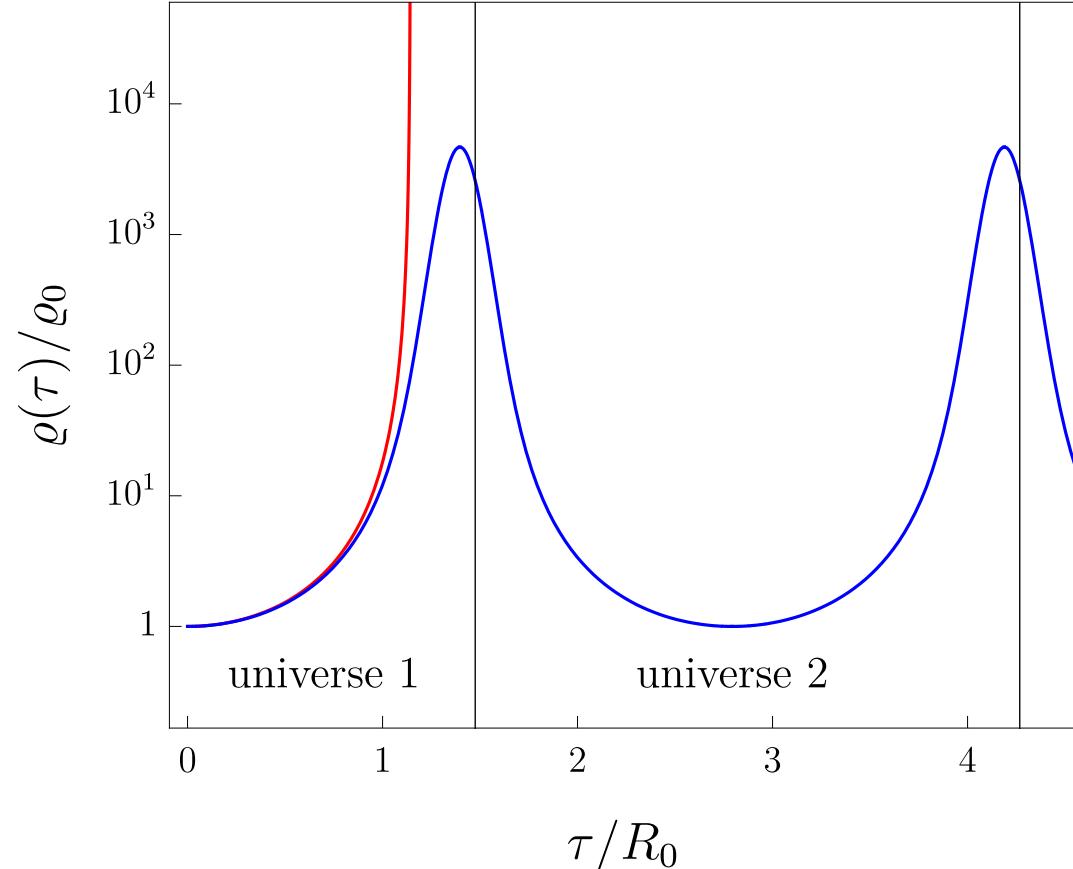
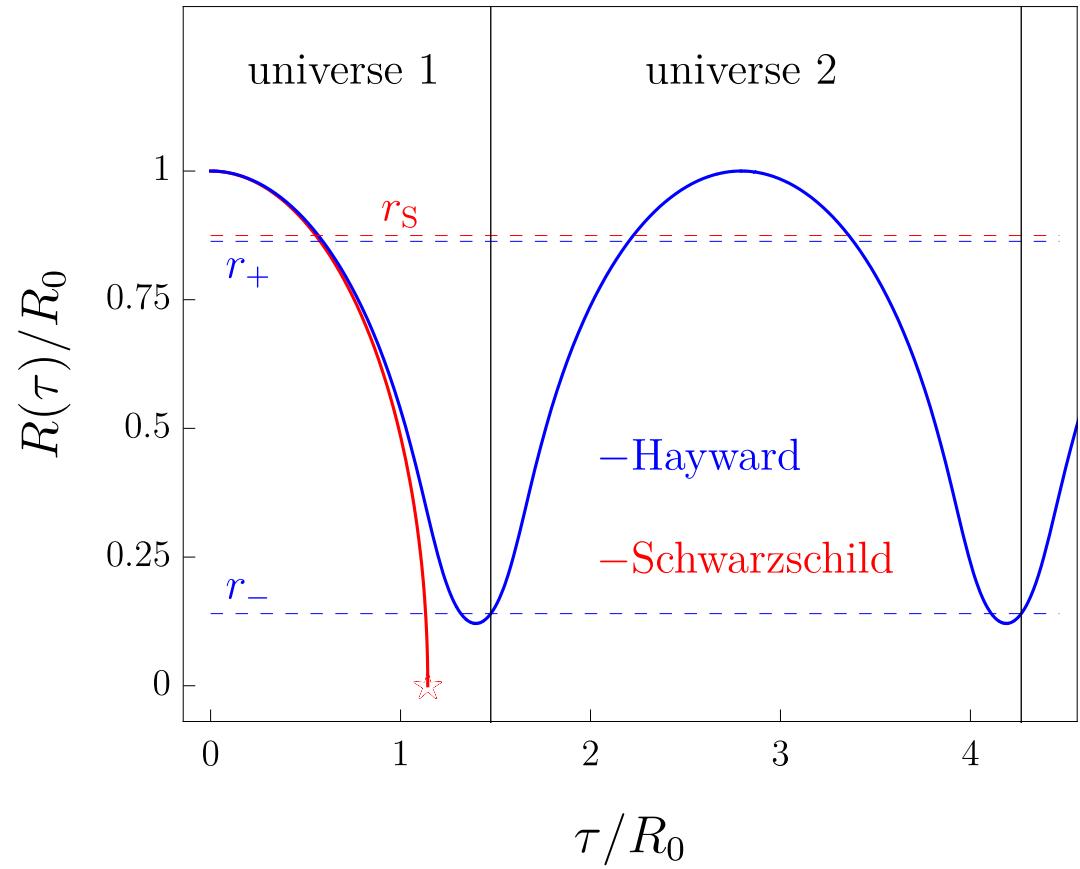
Spherical Symmetric Configurations of Matter

Oppenheimer-Snyder Dust Collapse

[Bueno, Cano, Hennigar, Murcia, AVC '25]

- ⊕ Equation for the star's radius R

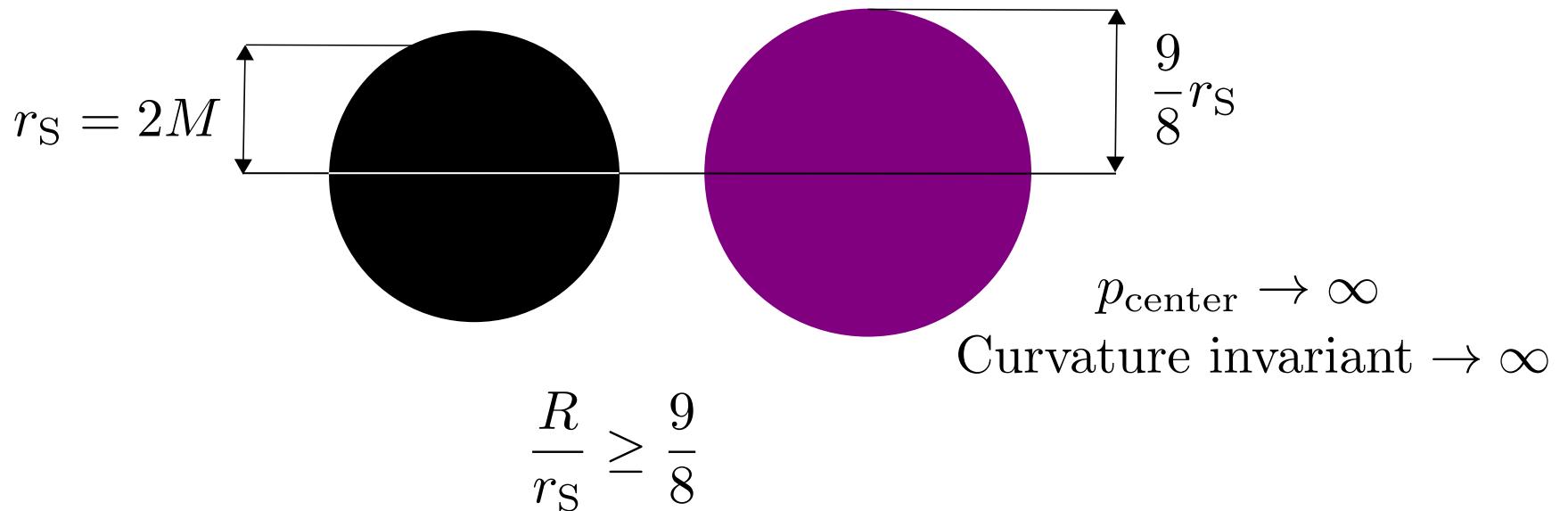
$$\dot{R}^2(\tau) + f(R(\tau)) = f(R_0)$$



Static Stars: Buchdahl Limit

[Buchdahl '59]

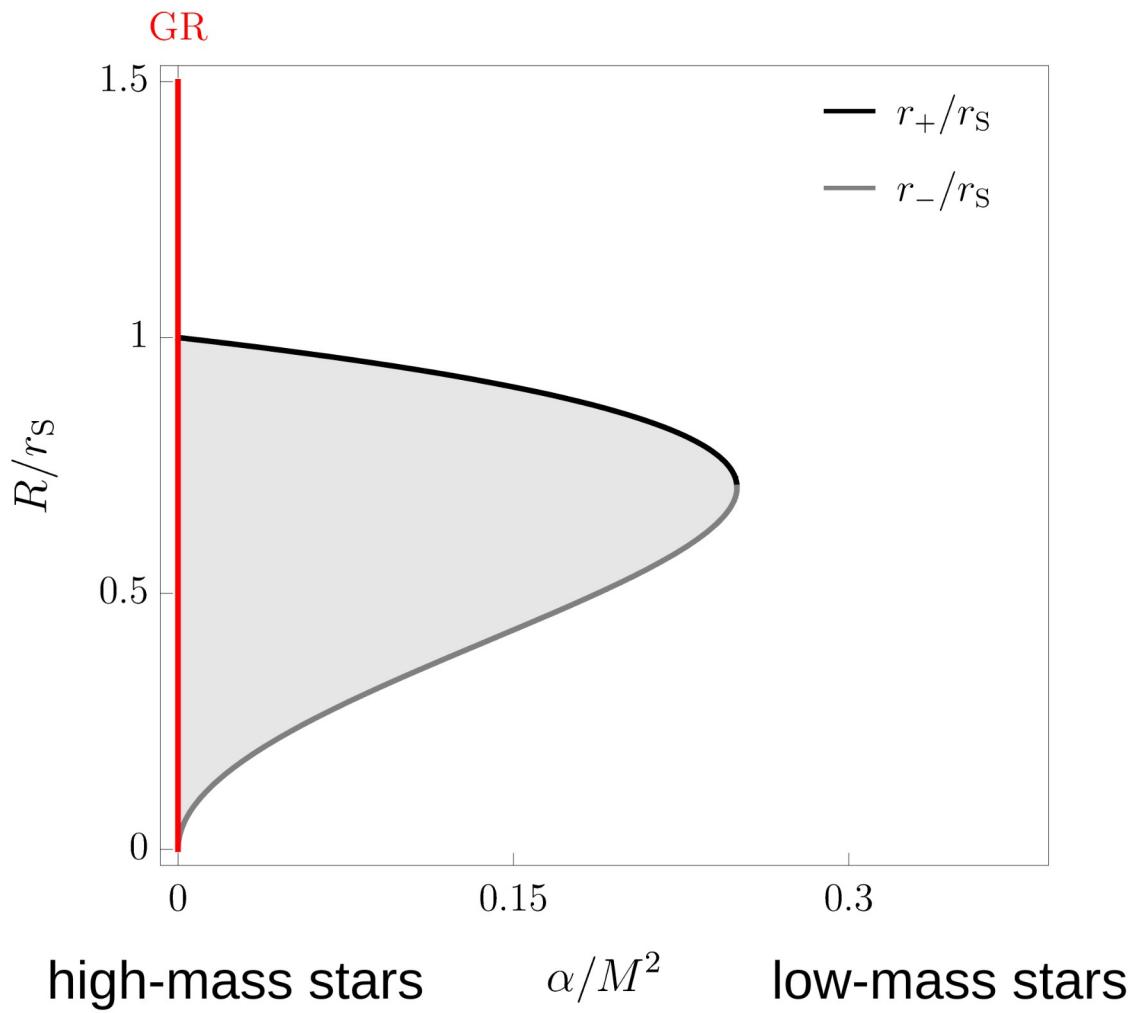
- ⊕ Upper bound on a star's compactness



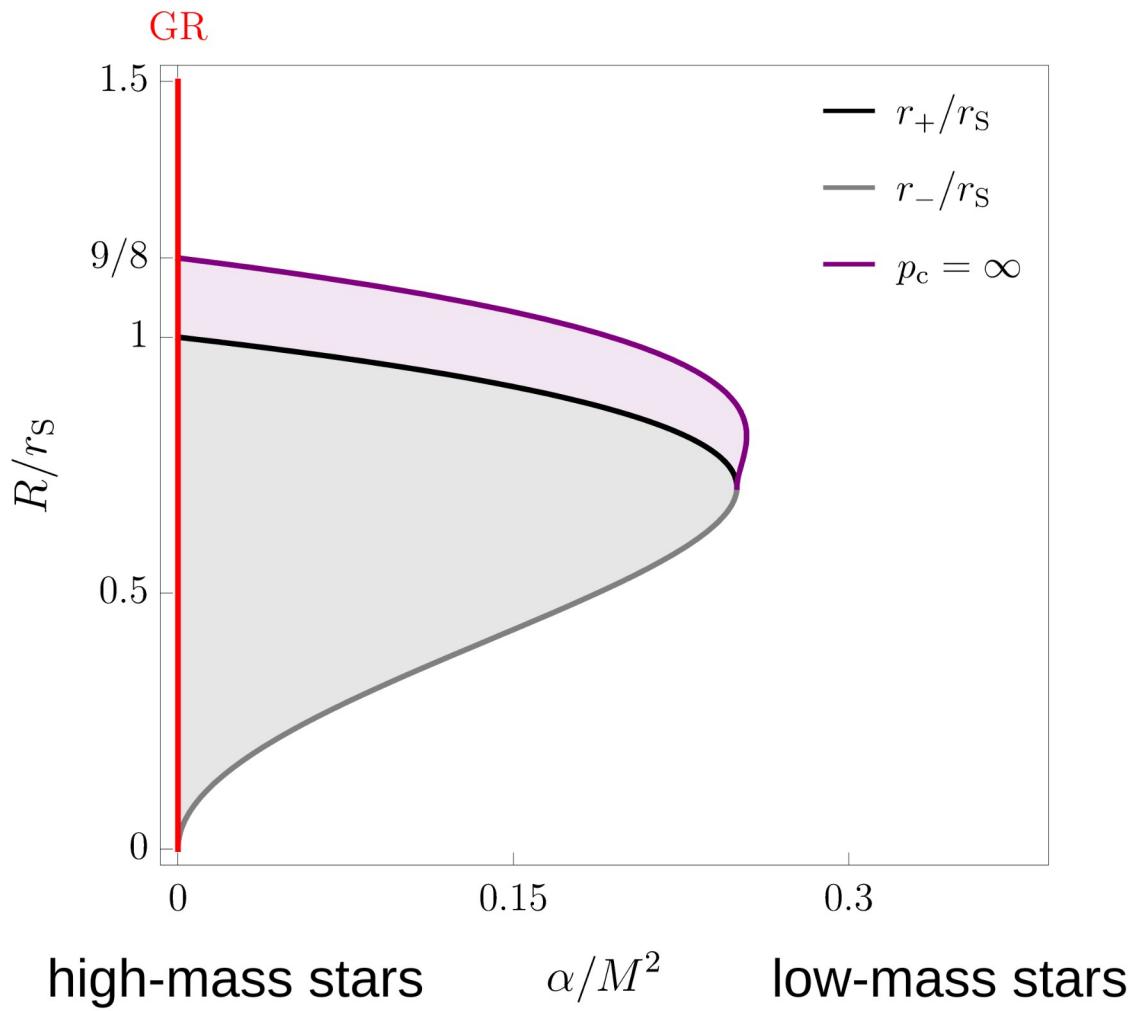
[Bueno, Hennigar, Murcia, AVC '25]

Buchdahl Limit for “Hayward” Stars with Constant-Density

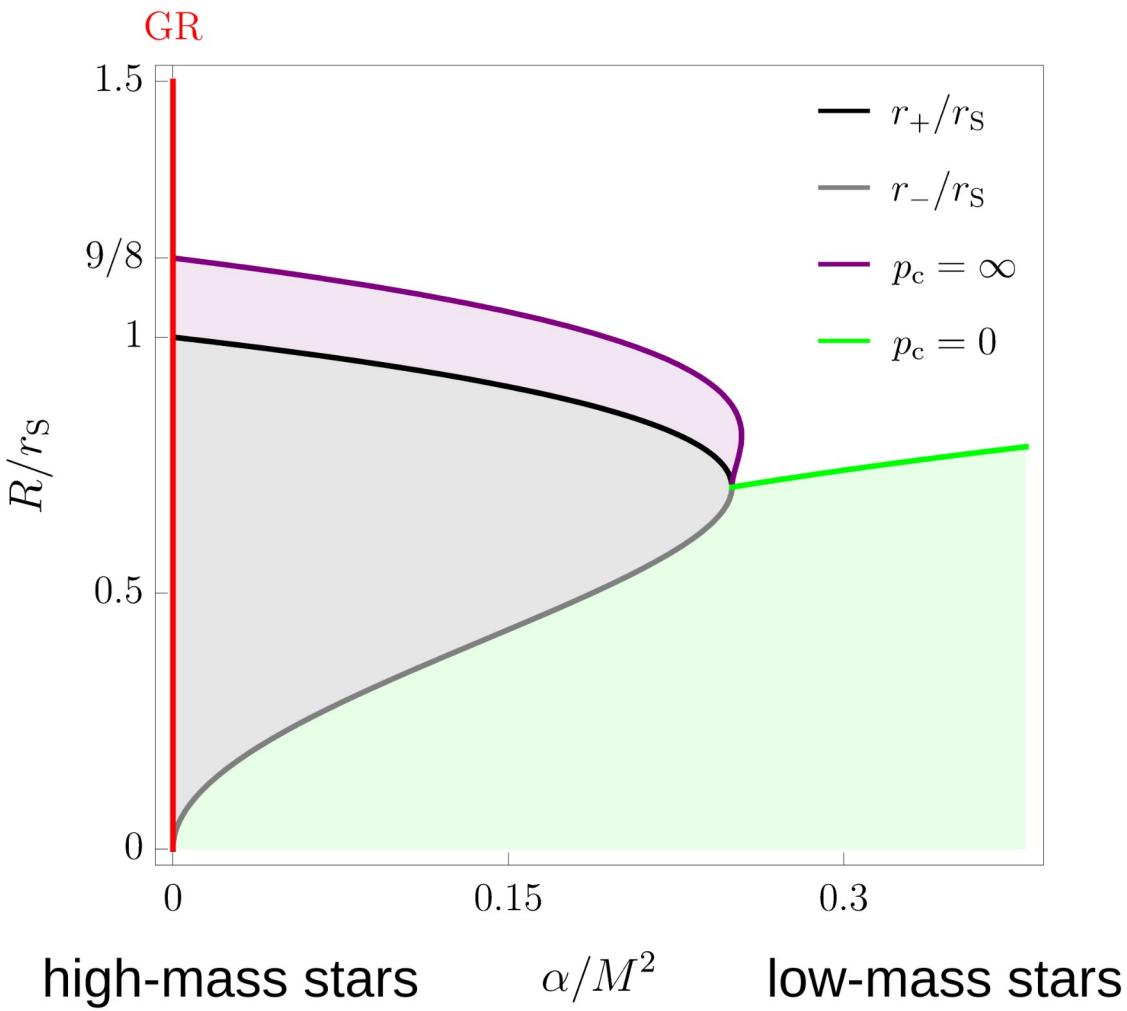
✿ Gray: Black Hole interior



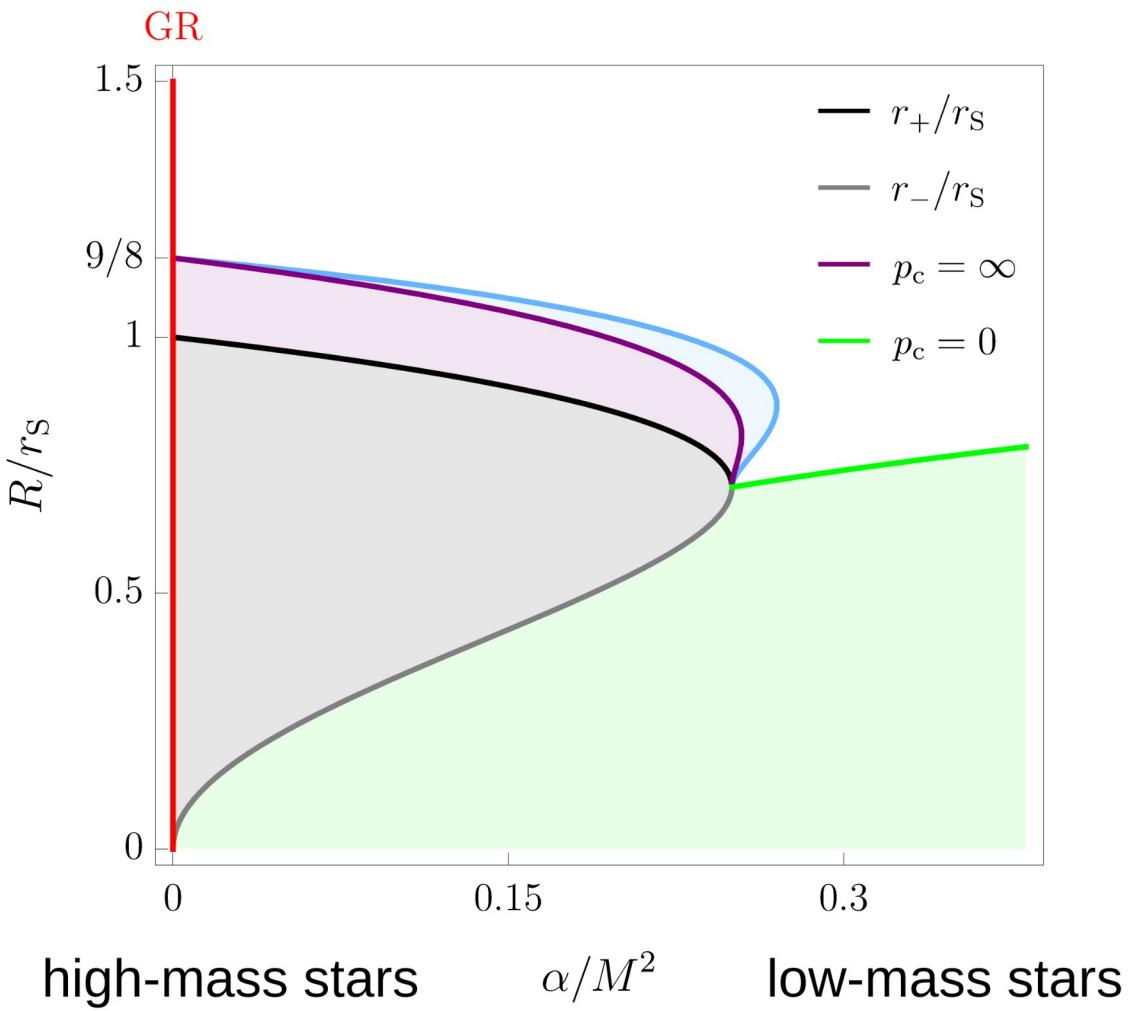
- ✿ Gray: Black Hole interior
- ✿ Purple: unphysical stars



- ⊕ Gray: Black Hole interior
- ⊕ Purple: unphysical stars
- ⊕ Green: exotic stars $p < 0$



- ✚ Gray: Black Hole interior
- ✚ Purple: unphysical stars
- ✚ Green: exotic stars $p < 0$
- ✚ Blue: high-curvature stars



Conclusions

- ⊕ Regular black holes as vacuum solutions of modified gravity.
- ⊕ Oppenheimer-Snyder collapse becomes an endless cycle of contraction and expansion.
- ⊕ Stars can reach higher compactness than in General Relativity.

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THANK YOU!