

From Stars to Regular Black Holes

Aitor Vicente-Cano

Based on 2505.09680 & 2512.19796

with Pablo Bueno, Pablo A. Cano, Robie A. Hennigar and Ángel J. Murcia

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Institut de Ciències del Cosmos
UNIVERSITAT DE BARCELONA



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Regular Black Holes from Pure Gravity

[Bueno, Cano, Hennigar '24]

$$S = \int \left(\mathcal{L}_{\text{GR}} + \sum_{n=2}^{\infty} \alpha_n \mathcal{L}_{(n)} \right)$$

Regular Black Holes from Pure Gravity

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$$ds^2 = -N(r)^2 f(r) dt^2 + \frac{dr^2}{f(r)} + r^2 d\Omega^2$$

A Regular Black Hole: Hayward

[Hayward '05]

$$\alpha_n = \alpha^{n-1}$$

A Regular Black Hole: Hayward

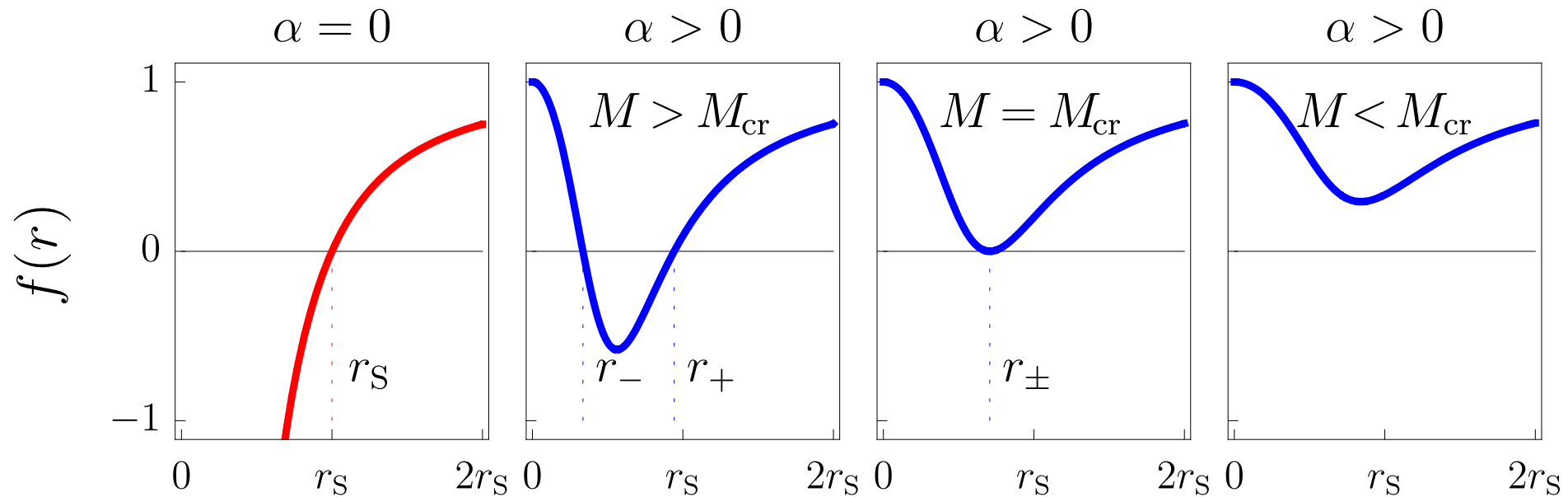
[Hayward '05]

$$\alpha_n = \alpha^{n-1} \quad \Rightarrow \quad f(r) = 1 - \frac{2Mr^2}{r^{D-1} + 2M\alpha}$$

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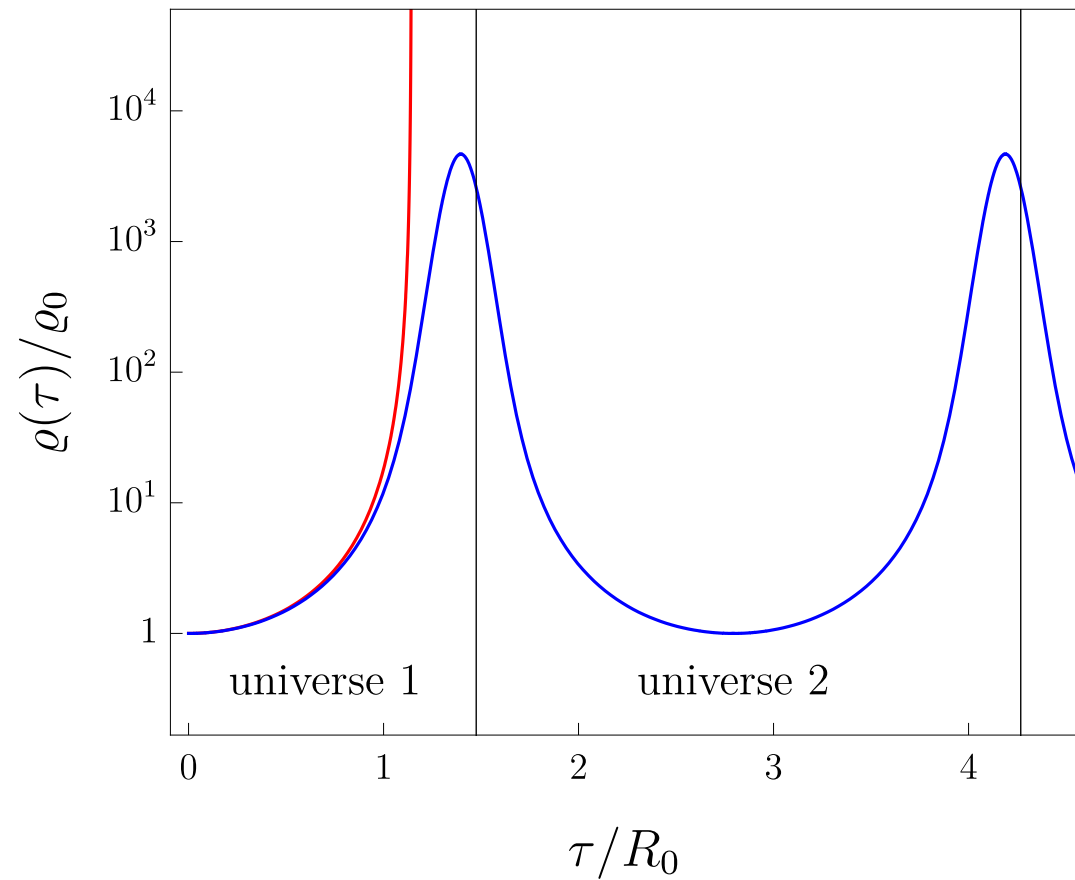
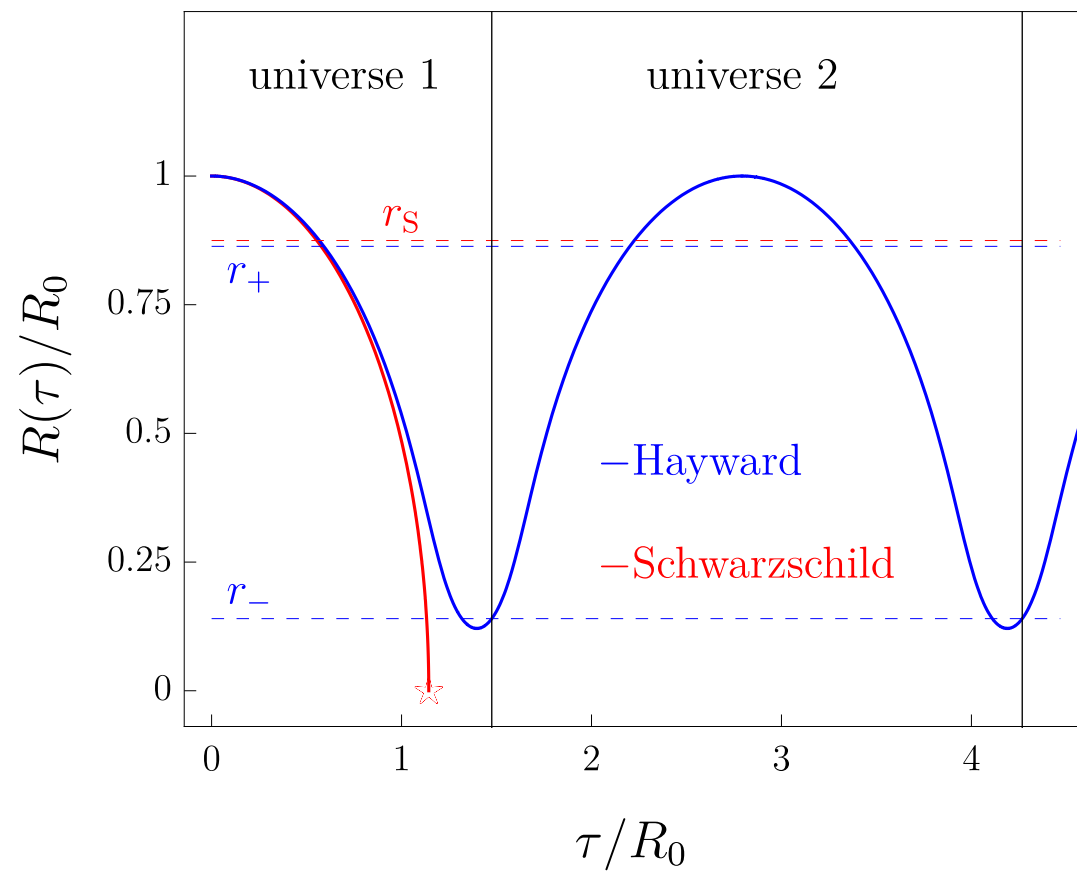
Spherical Symmetric Configurations of Matter

Oppenheimer-Snyder Dust Collapse

[Bueno, Cano, Hennigar, Murcia, AVC '25]

✦ Equation for the star's radius R

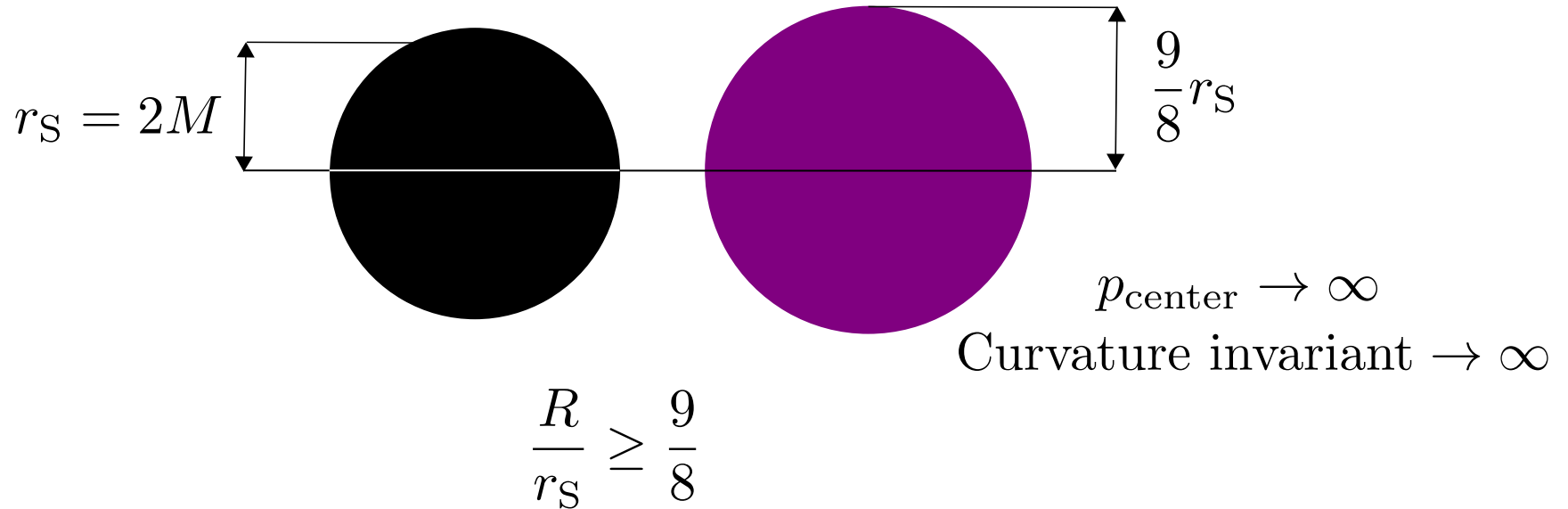
$$\dot{R}^2(\tau) + f(R(\tau)) = f(R_0)$$



Static Stars: Buchdahl Limit

[Buchdahl '59]

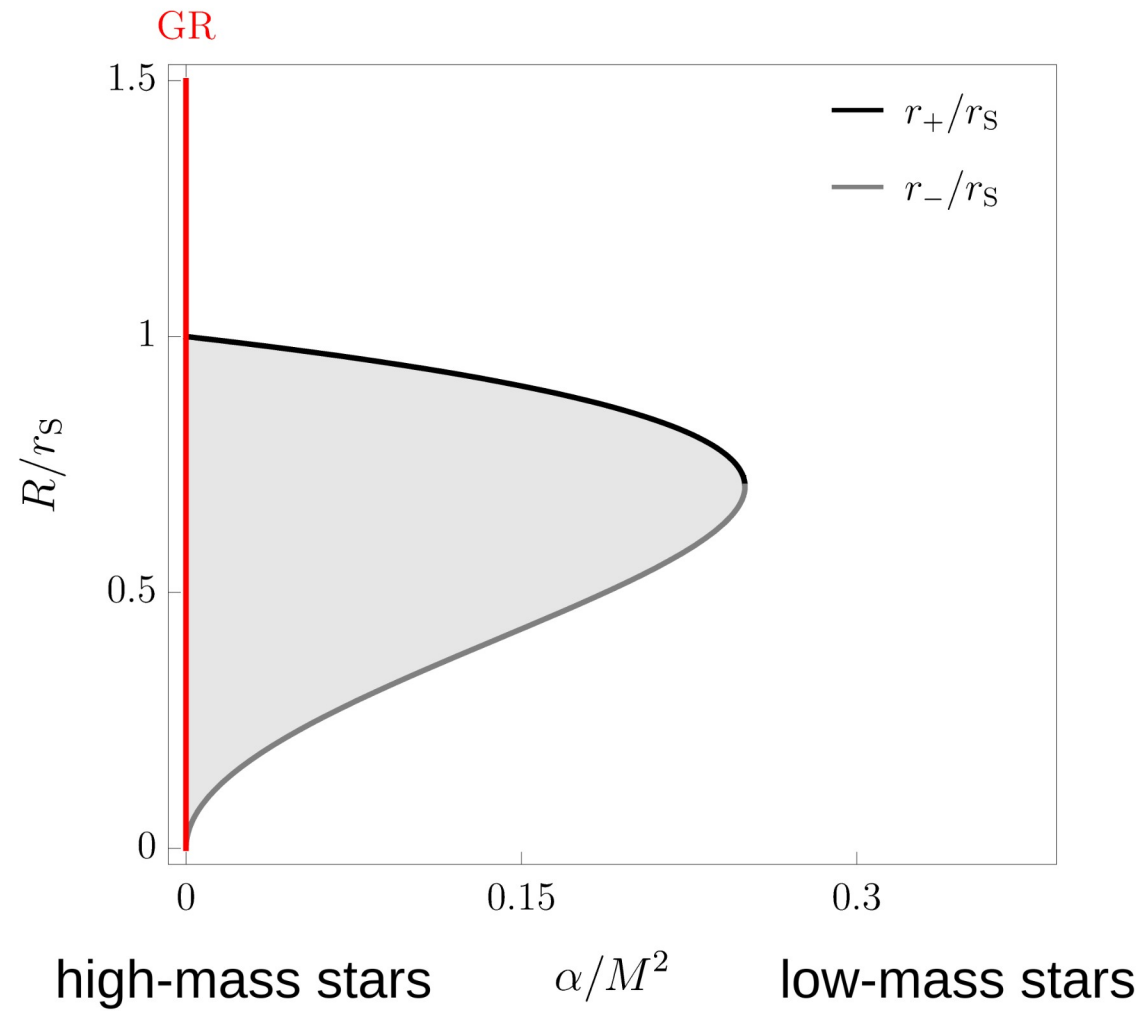
- ✦ Upper bound on a star's compactness



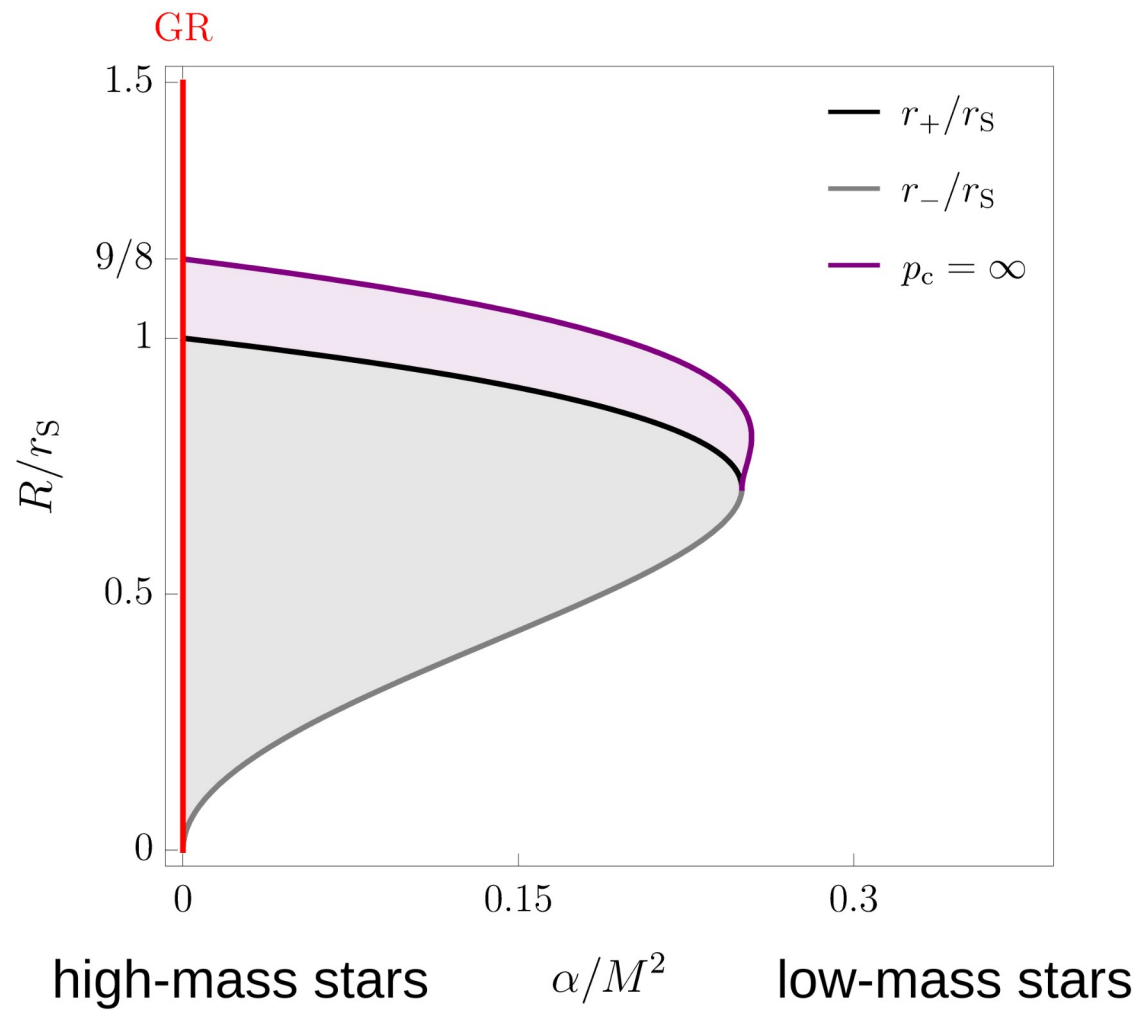
[Bueno, Hennigar, Murcia, AVC '25]

**Buchdahl Limit
for “Hayward” Stars
with Constant-Density**

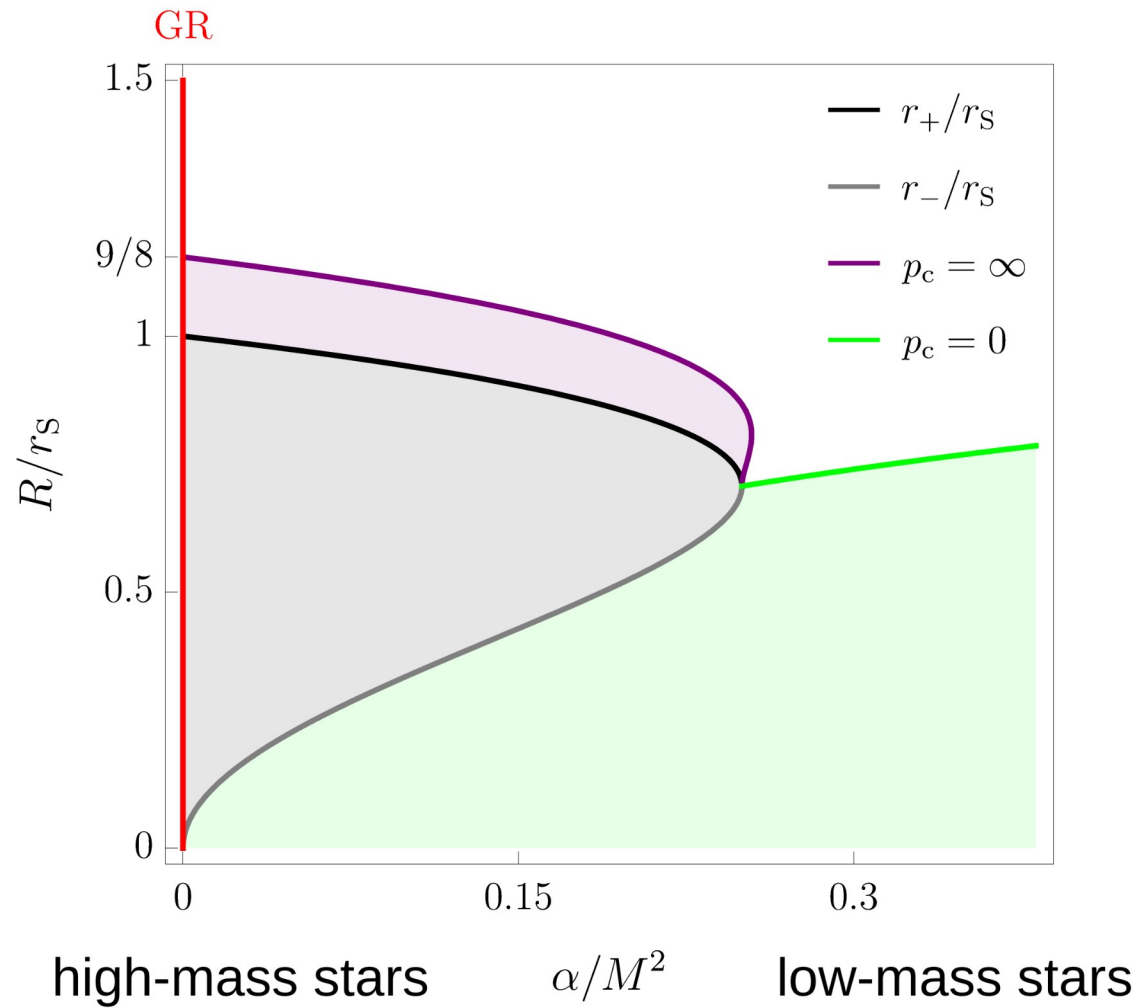
✦ Gray: Black Hole interior



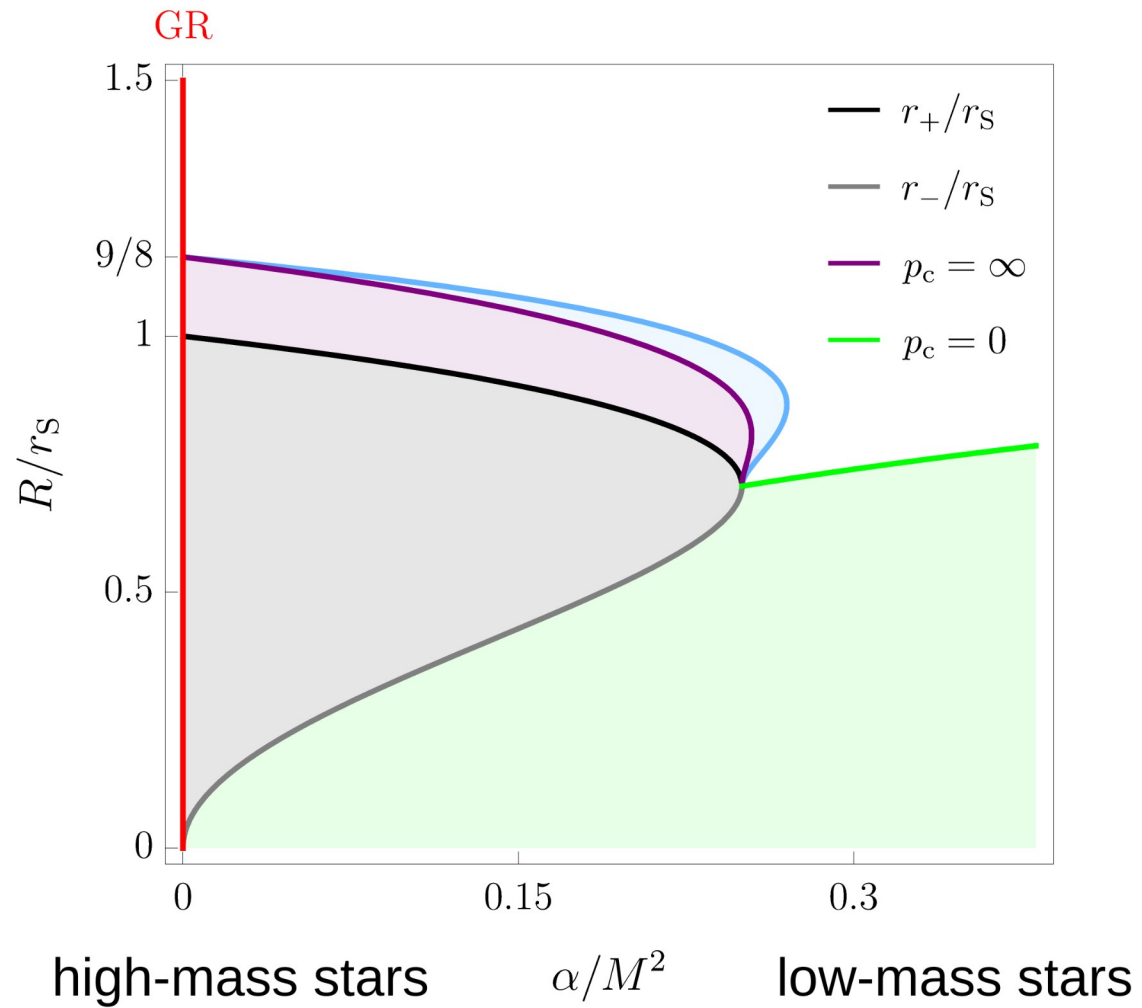
- ✦ Gray: Black Hole interior
- ✦ Purple: unphysical stars



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- ✦ Purple: unphysical stars
- ✦ Green: exotic stars $p < 0$



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- ✦ Purple: unphysical stars
- ✦ Green: exotic stars $p < 0$
- ✦ Blue: high-curvature stars



Conclusions

- ✦ Regular black holes as vacuum solutions of modified gravity.
- ✦ Oppenheimer-Snyder collapse becomes an endless cycle of contraction and expansion.
- ✦ Stars can reach higher compactness than in General Relativity.

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THANK YOU!