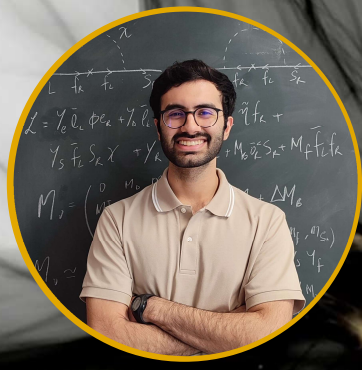


# The Dark Side of Neutrinos



## ADITYA BATRA

PhD Student @ CFTP (2023-2027)



Centro de Física Teórica de Partículas (CFTP)

FCT PhD Grant: UI/BD/154391/2023  
aditya.batra@tecnico.ulisboa.pt

**PhD Programme:** Neutrinos: a window to the Universe

**Supervisors:**

Filipe Joaquim (CFTP/IST)  
Rahul Srivastava (IISER Bhopal)  
José W. F. Valle (IFIC, València)

## 2022: MSc in Physics

**MSc Thesis:**

$h \rightarrow \gamma\gamma$  as a Novel Probe for New Physics

**Supervisor:**

Rahul Srivastava



MSc Thesis

## Highlighted Publications:

• **Axion framework with color-mediated Dirac neutrino masses**  
A. Batra, H.B. Câmara, F.R. Joaquim, N. Nath, R. Srivastava, J.W.F. Valle  
Published in: Phys.Lett.B 868 (2025) 139629

• **Axion paradigm with color-mediated neutrino masses**  
A. Batra, H.B. Câmara, F.R. Joaquim, R. Srivastava, J.W.F. Valle  
Published in: Phys.Rev.Lett. 132 (2024) 5, 051801

• **Dark linear seesaw mechanism**  
A. Batra, H.B. Câmara, F.R. Joaquim  
Published in: Phys.Lett.B 843 (2023) 138012

• **Phenomenology of the simplest linear seesaw mechanism**  
A. Batra, P. Bharadwaj, S. Mandal, R. Srivastava, J.W.F. Valle  
Published in: JHEP 07 (2023) 221

•  **$h \rightarrow \gamma\gamma$  Decay: Smoking Gun Signature of Wrong-Sign  $hbb$  Coupling**

A. Batra, S. Mandal, R. Srivastava  
e-Print: 2209.01200 [hep-ph]



Full Publication List

## Talks at international conferences:

- **FLASY 2025** - 11th Workshop on Flavour Symmetries and Consequences in Accelerators and Cosmology (Roma Tre, Italy)
- **Corfu2024** - Workshop on Standard Model and Beyond (Corfu, Greece)
- **ISAPP 2024** - Particle Candidates for Dark Matter (SGGS, Padua, Italy)
- **PLANCK2024** - The 26th International Conference "From the Planck Scale to the Electroweak Scale" (IST, Lisboa, Portugal)



UIDB/00777/2020,  
UIDP/00777/2020,  
CERN/FIS-PAR/0019/2021,  
UI/BD/154391/2023



**The problems:** the **Standard Model** cannot explain:

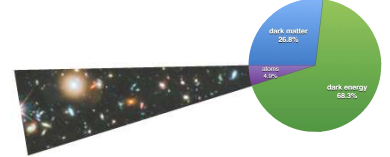
### Neutrino flavour oscillations

- Neutrinos can change from one type to the other while they propagate.
- This is only possible if they have non-zero masses.



### Observed Dark Matter abundance

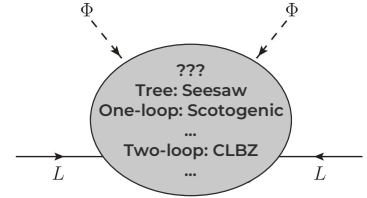
- Cosmological evidence suggests that 26.8% of the total matter in the Universe appears in the form of dark matter.



## Neutrino masses

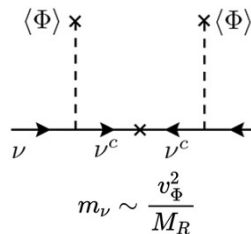
The unique lowest-dimensional effective operator that can be added to the Standard Model is the dimension-five **Weinberg operator**. This operator leads to **Majorana masses** for neutrinos after electroweak symmetry breaking.

$$-\mathcal{L}_{\text{Maj.}}^{d=5} = \frac{\kappa_{\text{Maj.}}}{\Lambda} (\bar{L}^c \tilde{\Phi}^*) (\tilde{\Phi}^\dagger L) + \text{H.c.}$$



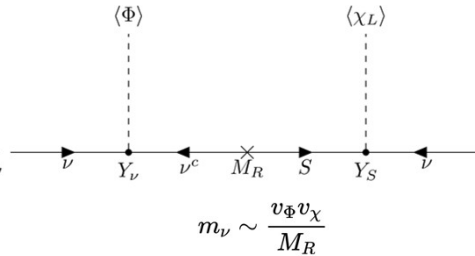
## Popular solutions

### Type-I Seesaw Model



$$m_\nu \sim \frac{v_\Phi^2}{M_R}$$

### Linear Seesaw Model



$$m_\nu \sim \frac{v_\Phi v_\chi}{M_R}$$

➤ The **Type-I Seesaw** is by far the **simplest solution** to the neutrino mass problem.

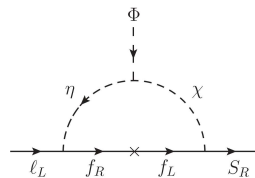
➤ A major drawback of this model is the large mass scale of the right-handed neutrinos, far away from the reach of current experiments.

➤ The **Linear Seesaw**, despite being more complicated, is a

**low-scale solution**

that offers more testability prospects at ongoing experiments.

## Our approach:

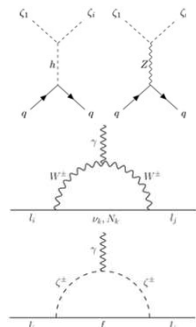


The **lepton number symmetry** is violated by the scalar potential term:

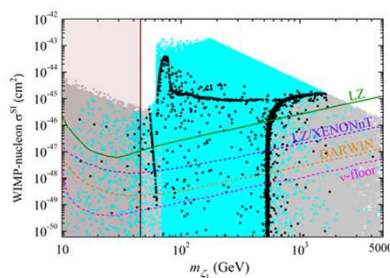
$$V_{\text{soft}} = \kappa (\eta^\dagger \Phi) \chi + \text{H.c.},$$

We propose a model where the **low-scale linear seesaw** neutrino mass generation mechanism is seeded by cosmologically stable dark matter particles accounting for both **neutrino flavour oscillations** and the observed **dark matter** abundance.

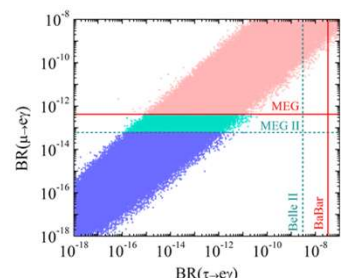
## The results



### Dark matter direct detection



### Charged lepton flavour violation



The **scalar dark matter** particles can interact with normal matter directly through the Higgs or Z boson. Furthermore, the new particles can mediate **charged lepton flavour violating decays** with sizable branching ratios. Therefore, our model can be probed through these processes at various current and upcoming experiments.