

Relativistic Plasma Mirrors

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Motivation

Attosecond pulses unlock access to electron dynamics and ultrafast phenomena across atomic, molecular, and plasma physics. A key challenge is generating them with sufficient intensity for applications. Density spikes at the rear of a laser-driven plasma bubble act as relativistic mirrors [1, 2], transferring momentum to a counter-propagating pulse via the double Doppler effect, compressing and upshifting it to higher intensities.

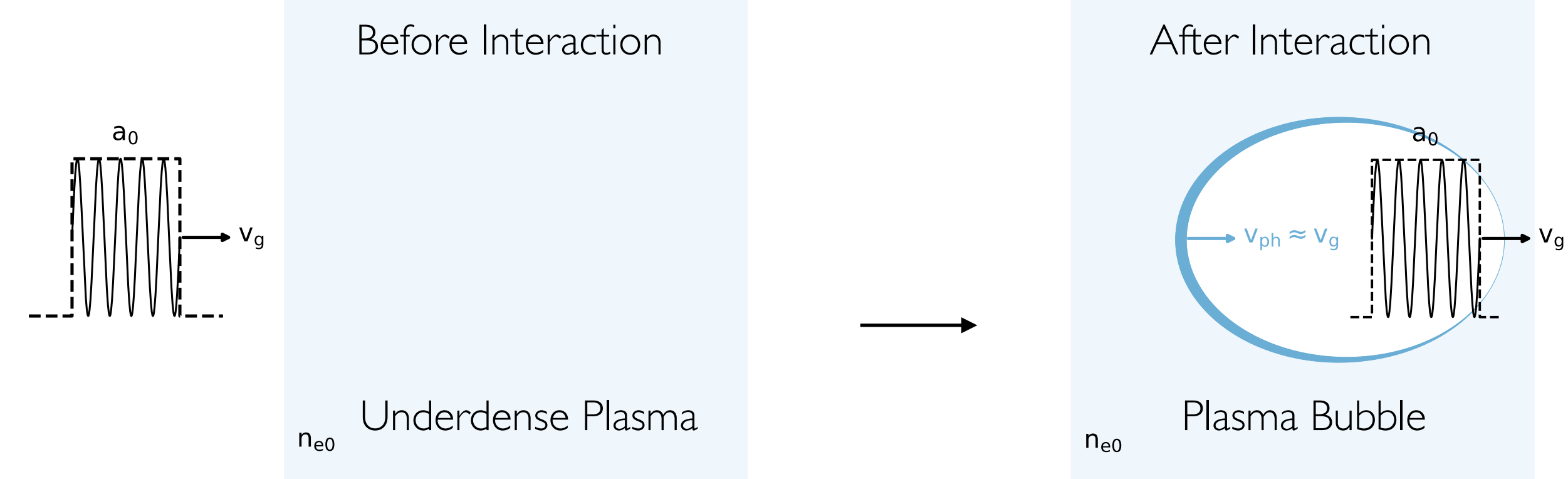


Figure 1: An intense laser driver propagates through an underdense plasma and forms a plasma bubble in its wake.

Density spike parameters at the rear of the plasma bubble

In the 1D quasi-static approximation for a cold, collisionless plasma:

$$n_e(\xi) = \frac{1}{2} \left[1 + \frac{1 + a^2(\xi)}{y^2(\xi)} \right]$$

$$\gamma_e(\xi) = \frac{1}{2} \left[y(\xi) + \frac{1 + a^2(\xi)}{y(\xi)} \right]$$

The maximum density, Lorentz factor, and thickness of the plasma mirror are [3]:

$$n_{e \max} = \frac{1}{2} (1 + \gamma_{\perp}^4) \quad \gamma_{\max} = \frac{n_{e \max}}{\gamma_{\perp}^2}$$

$$l_{\text{mirror}} = 4\gamma_{\perp} \text{Im} \left[E \left(\arcsin \gamma_{\perp} \middle| \frac{1}{\gamma_{\perp}^2} \right) - F \left(\arcsin \gamma_{\perp} \middle| \frac{1}{\gamma_{\perp}^2} \right) \right]$$

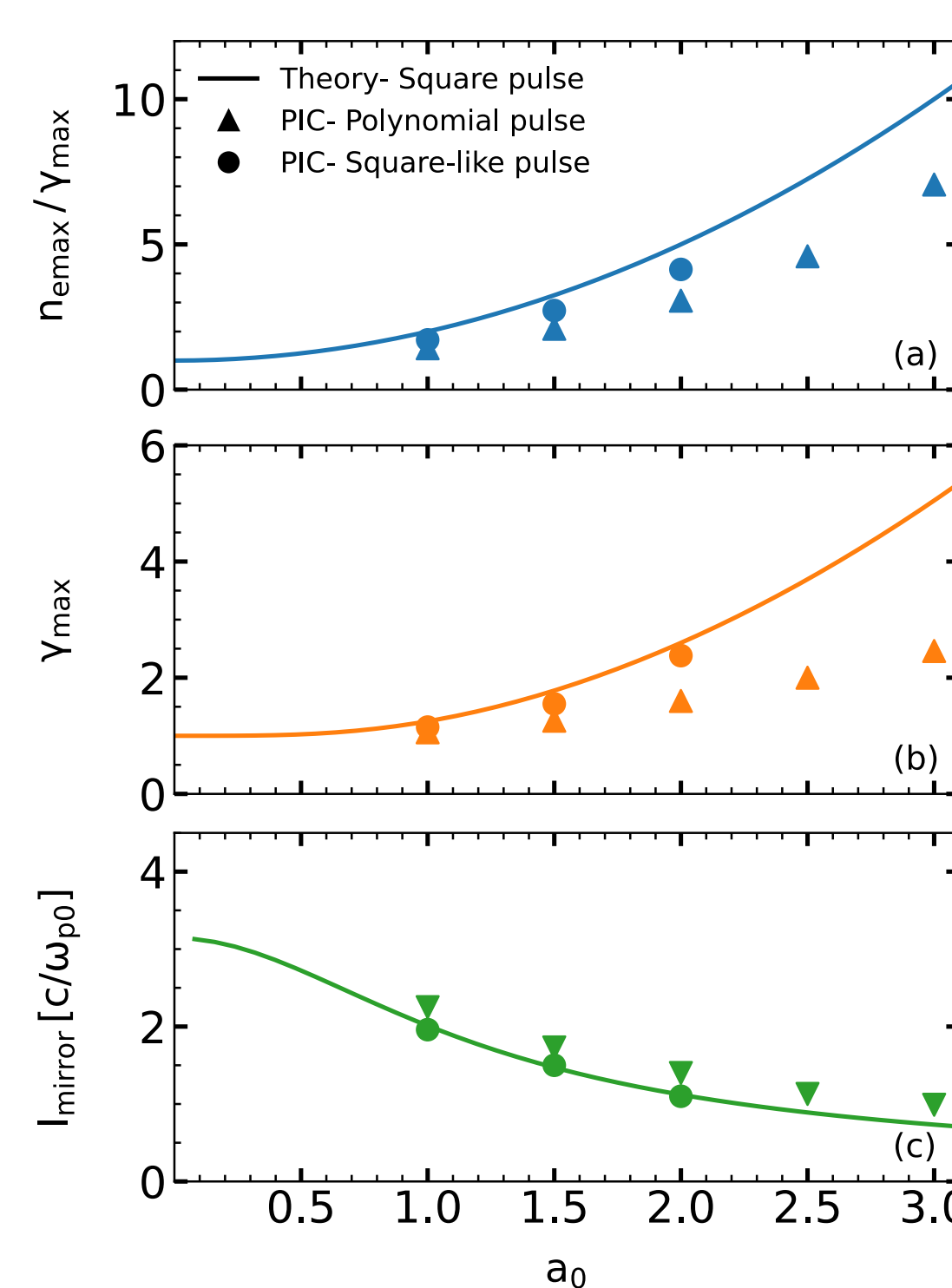


Figure 2: OSIRIS [4] results follow the analytical trends; simulation parameters are consistent with bubble regime theory [5] and modern laser facilities.

Density profile of the plasma mirror

Osiris

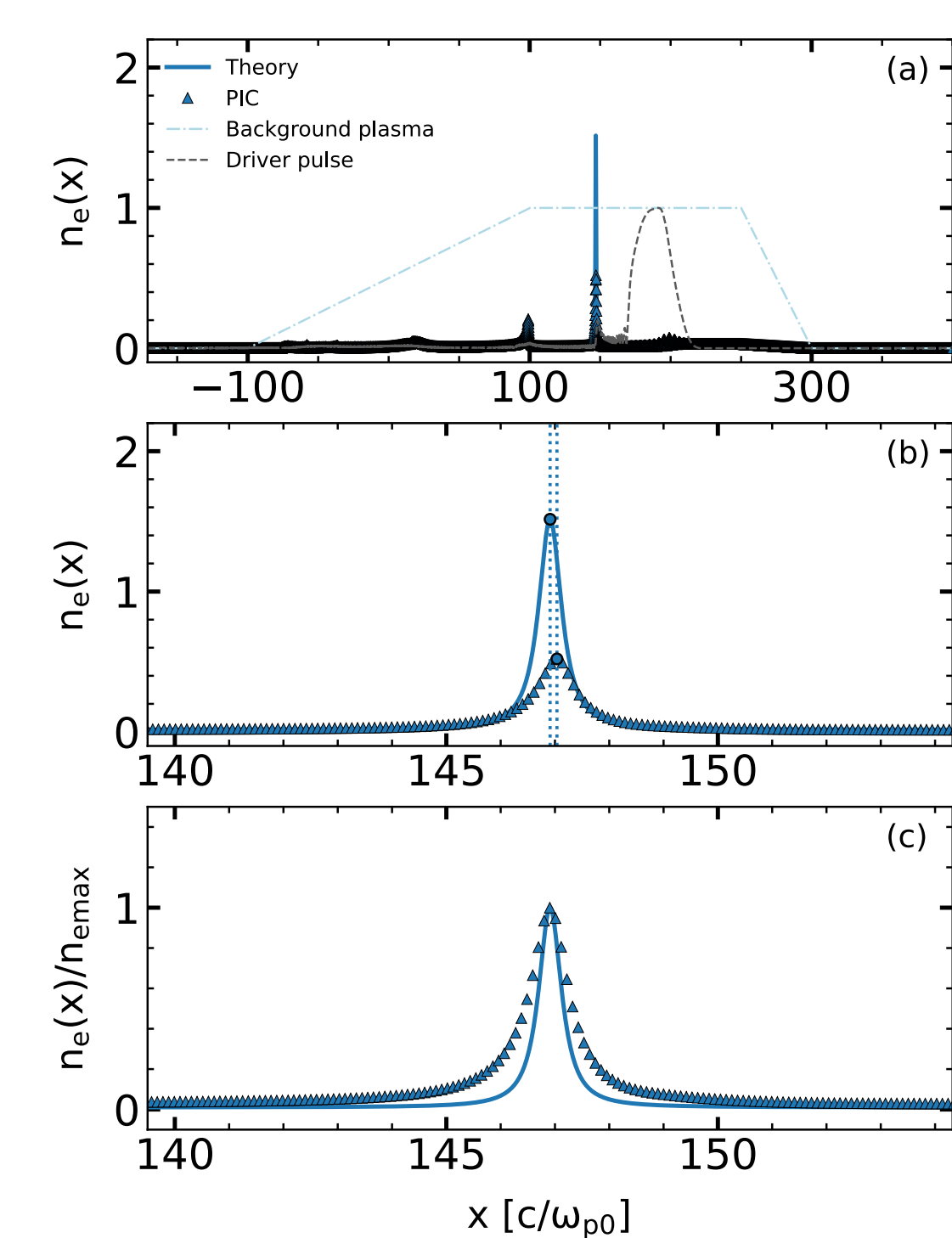


Figure 3: OSIRIS data validate the analytical density profile model.

Expanding the density spike about its maximum shows it is locally parabolic [3].

$$n_e(\xi) = [(1 + \gamma_{\perp}^4)/2] - [(\gamma_{\perp}^6(\gamma_{\perp}^4 - 1)/4)(\xi - \xi_{\text{peak}})^2] + O[(\xi - \xi_{\text{peak}})^4]$$

Reflection coefficient for flat-top, parabolic, and Dirac-limit plasma mirrors

Solving the Helmholtz equation in the mirror rest frame for three density profiles and transforming to the laboratory frame yields:

$$|R_{\text{flat}}| = \Gamma^2 \left| \frac{\omega_{\text{cut-off}}^2 - \frac{\omega_{p0}^2}{\Gamma^2}}{(2\omega_0^2 - \omega_{\text{cut-off}}^2 - \frac{\omega_{p0}^2}{\Gamma^2}) + 2i\chi\eta\sqrt{1 - \frac{\omega_{p0}^2}{\Gamma^2\omega_0^2}}\omega_0^2 \cot(\chi\eta\gamma^2(1 + \eta\beta)k_0 l_{\text{mirror}})} \right|$$

$$|R_{\text{Dirac}}| = \Gamma^2 \left| \frac{1 - \frac{\omega_{p0}^2}{\omega_{p \max}^2}}{1 + \frac{\omega_{p0}^2}{\omega_{p \max}^2} - \frac{2i\sqrt{1 - \frac{\omega_{p0}^2}{\Gamma^2\omega_0^2}}(1 + \eta\beta)\omega_0^2}{k_0\omega_{p \max}^2}} \right|$$

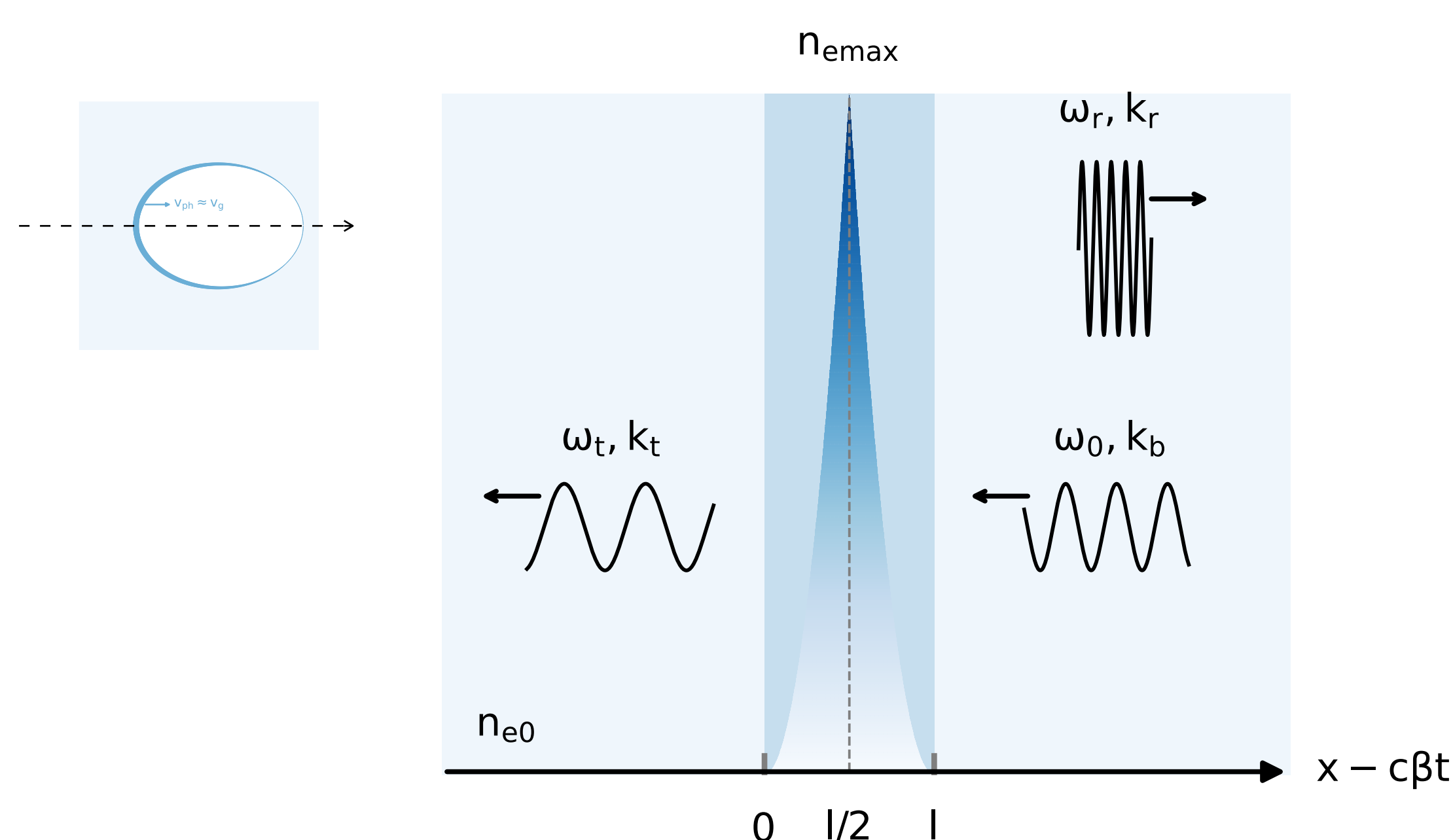


Figure 4: Laboratory-frame configuration of the moving mirror for flat-top, parabolic, and Dirac-limit density profiles.

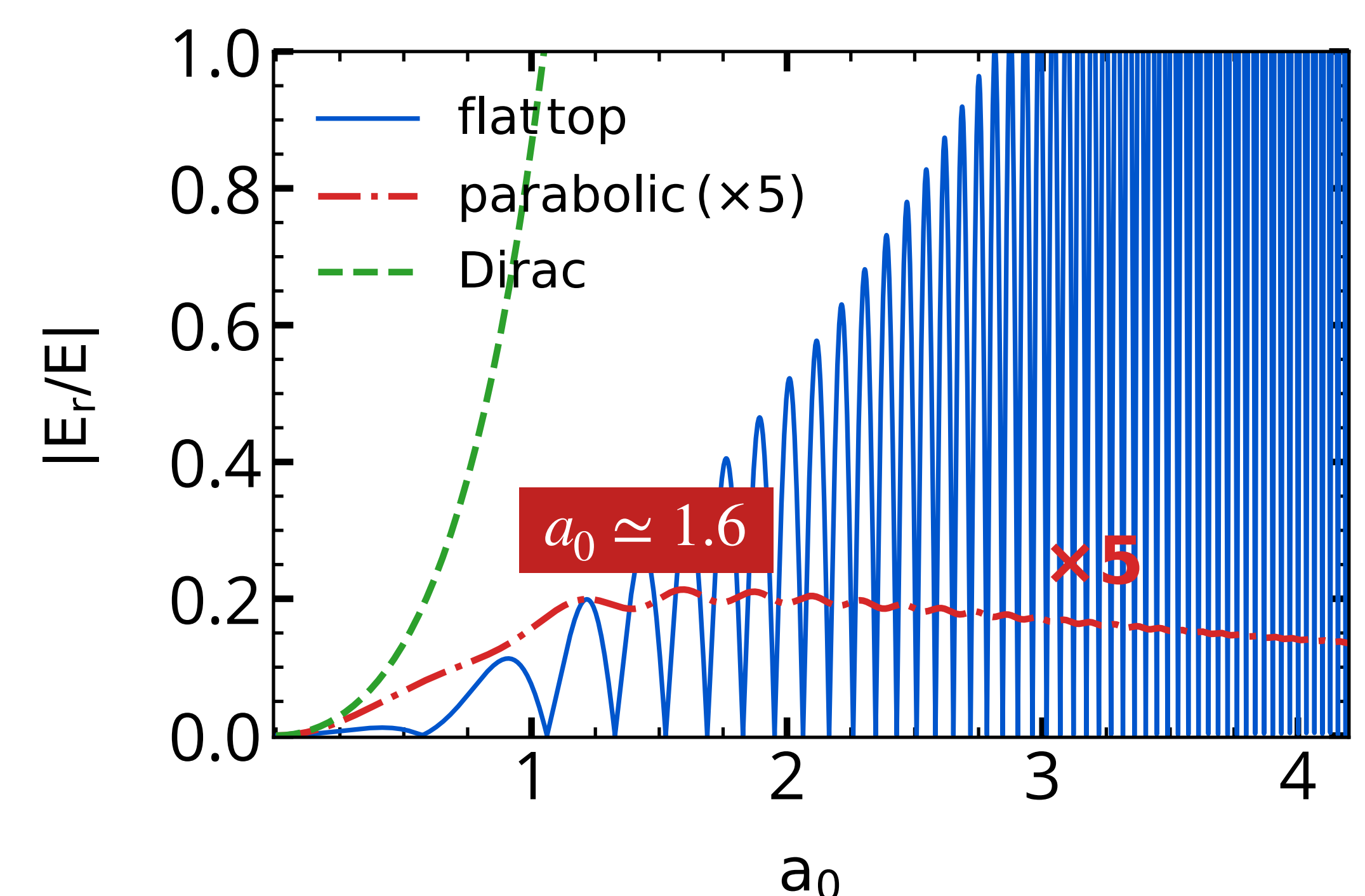


Figure 5: The magnitude of the reflection coefficient, $|R| = |E_r/E|$, versus a_0 for flat-top, parabolic, and Dirac-limit density profiles ($\omega_0 = 2\omega_{p0}$).

Conclusions

Real plasma mirrors have finite thickness and a non-ideal density profile.

Under the quasi-static approximation, $|R|$ can be expressed as a function of a_0 alone (for fixed ω_0/ω_{p0}).

2D OSIRIS simulations confirm the density spike is well described by a parabolic profile.

The parabolic mirror reflectivity peaks at an optimal a_0 , set by the balance between electron compression, mirror thickness, and relativistic transparency.

References

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